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## Stau Pair Production At The ILC

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### Outline

• Introduction of the International Linear Collider (ILC)

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- Introduction of  $(g_{\mu}-2)$  in the MSSM
- Motivation
- Stau and Neutralino Reconstruction
- Results
- Summary
- Future Plans

### International Linear Collider (ILC)

e<sup>+</sup>

#### High Energy e<sup>+</sup> e<sup>-</sup> Linear Collider:

- Clean background (less synchrotron radiation)
- Able to polarize electron and positron. ( $P(e^-, e^+)=(0.8, -0.3)$ )
- Measure the mass, spin and interaction strength of the Higgs boson.
- Able to search for extra dimension and SUSY if exist.
- Search for Neutralino (possible candidate for cold dark matter)

### How good will the ILC be?

g(hAA)

 $g(hAA)|_{SM}$ 



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[Ref]: arXiv:1207.2516v3 [hep-ph] - M.Peskin

### New Physics Beyond The Standard Model

### Incomplete of the Standard Model:

- The hierarchy problem
- Gravity
- Matter/antimatter asymmetry.
- Dark Matter (weakly interacting massive particle WIMP)



27% Matter = 4% baryonic + 23% non-baryonic



At the ILC: MSSM, SUGRA, Extra Dimensions could be test.

### SUSY

- At TeV scale,
  - SUSY is the resolution of the hierarchy problem in the SM.
  - The unification of couplings (MSSM)
  - The possible candidate for dark matter
  - Low-energy SUSY is strongly suggested by string theory



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# $(g_{\mu} - 2)$ in the MSSM

• The anomalous magnetic moment in  $(g_{\mu}-2)$ :  $a_{\mu} \equiv (g_{\mu}-2)/2 \qquad g_{\mu}=2 \text{ (tree)} \\ g_{\mu}\neq 2 \text{ (RC)}$ 

$$\Delta a_{\mu} = a_{\mu}(\exp) - a_{\mu}(SM) = (26.1 \pm 8.0) \times 10^{-10}$$

The muon g-2 anomaly indicates physics BSM at more than 3σ
 [Ref]. "Muon g-2 vs LHC in SUSY Models" – M. Endo, K. Hamaguchi

- In MSSM, it is very difficult to get a Higgs boson mass as large as 115 GeV  $\Rightarrow$  The correlation between the Higgs mass and  $g_{\mu}$ -2 is significant.
- MSSM contribution to the  $g_{\mu}$ -2 are enhanced by tan $\beta$ =  $\langle H_{u} \rangle / \langle H_{d} \rangle$  = O(10)
- Non color super-partners ( slepton, neutralinos and charginos ) are light due to the  $g_{\mu}$ -2 discrepancy.

 $M_{SUSY} = O(100 \text{ GeV}) \text{ for } \tan\beta = O(10)$ 

## Motivation for Stau pair production

- R-parity:  $R = (-1)^{3B+L+2S}$ 
  - The super-partners have R = -1
  - The SM particles have R = +1

- B: Baryon number L: Lepton number S: Spin
- There is no interaction coupling a single super-partner with two SM particles

The lightest SUSY particle (LSP) is absolutely stable O(100 GeV $\sim$ 1 TeV), interact weakly (WIMP). The next lightest SUSY particle (NLSP) is believed that is Stau

These two masses could be reached in ILC's range! ( $\sqrt{s}=250, 350, 500 \text{ GeV}$ )

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### Benchmarks for SUSY searches



[Ref]: "The Snowmass Points and Slopes: Benchmarks for SUSY Searchs" - arVix: hep-ph/0202233v1

### Method

(Using Physsim and Whizard to create the cross-section for Stau) In collider experiment LSP will carry away P<sup>µ</sup> while escaping detector

Super-partners → LSP+SM particle

Calculating tau decay modes using Tauola

[Ref]: "Long life stau in the MSSM" – T. Jittoh, J.Sato

e+

2-body decay mode



3-body decay mode



4-body decay mode  $v_{\tau}$   $v_{\tau}$   $v_{\tau}$  W

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 $Z^*/\gamma$ 

### Study at neutralino mass = 300 GeV



### Stau Search At The ILC

 $1^{st}$  layer (Radius = 16 mm)

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Depend on the life-time of Stau (Leptonic or Hadronic decay)

Impact parameter `

### Simulation

Signal:  

$$e^+e^- \rightarrow \tilde{\tau}\bar{\tilde{\tau}} \rightarrow \tau^+\tau^-\tilde{\chi}\tilde{\chi}$$



#### Back ground:

 $e^+e^- \rightarrow WW \rightarrow llvv$   $e^+e^- \rightarrow ZZ \rightarrow llvv$   $e^+e^- \rightarrow Z^*/\gamma^* \rightarrow ll$   $e^+e^- \rightarrow \tau\tau$  $e^+e^- \rightarrow bhabha$ 

#### Set up:

 $\sqrt{s} = 250, 350, 500 \text{ GeV}$ Beam polarization: P(e<sup>-</sup>,e<sup>+</sup>)=(0.8, -0.3) Stau mass = 120 ~ 240 GeV  $\theta_{\tau} = 0 ~ \pi$  rad Neutralino mass = 100 GeV Luminosity = 250 fb<sup>-1</sup>, 350 fb<sup>-1</sup>, 500 fb<sup>-1</sup> tan $\beta$ = 5, 10 -13-

### Polarization And Cross-section



### Physsim Simulation(1)



Generating cross-section of Stau: Check!

S-channel liked processes = 0K!

Figure1: Relation of cross-section and central mass energy corresponding to different stau masses. (Mixing angle is set at 0) Black: M-stau = 120 GeV Red: M-stau = 150 GeV Blue: M-stau = 200 GeV

### Physsim Simulation(2)



$$\begin{pmatrix} \tilde{\tau}_1 \\ \tilde{\tau}_2 \end{pmatrix} = \begin{pmatrix} \cos\theta_{\tau} & -\sin\theta_{\tau} \\ \sin\theta_{\tau} & \cos\theta_{\tau} \end{pmatrix} \begin{pmatrix} \tilde{\tau}_L \\ \tilde{\tau}_R \end{pmatrix}$$

**Figure3**: Relation of cross-section and stau mass (Ecm = 500 GeV,  $\theta_{\tau}$ = 0)

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### Whizard Simulation(1)

- Whizard provides more variation on parameters
  - Difficult to handle many parameters simultaneously
  - More study is required in order to be able to determine the parameters

Studying Susy-hit for testing muon g-2 in the MSSM

•  $M_1 = 90.8 \text{ GeV}$  (Bino  $\approx$  Neutralino)

- M<sub>2</sub> = μ= 1000 GeV (supposed to be unobservable – heavier than Neutralinos and Charginos)
- $\Gamma(\text{stau1}) = 1.97956261 \times 10^{-02} \text{ GeV}$
- $\Gamma(\text{stau2}) = 7.34654619 \times 10^{-01} \text{ GeV}$
- Branching ratio
  - BR(stau1 $_{10}$  $\tau^{-}$ ) = 100%
  - BR(stau2 $_{10}$   $\tau^{-}$ ) = 85.8%
  - BR(stau2  $\rightarrow$  snu\_tauL W<sup>-</sup>) = 12.9%
  - BR(snu\_tauL  $\rightarrow \chi_{10}$  nu\_tau) = 100%

(After discussing with Motoi Endo San)

### Whizard Simulation(2)

(Simulation using stau1stau1\_tau-tau process in MSSM model)

Stau-width for 2-body decay dominance  $(\delta m > m_{\tau})$ :



The tan $\beta$ = 6.1 and sin $\theta_{W}$  = 0.23116 (LEP) are set up for muon g-2 in the MSSM

Changing width leads to the change of Stau-mass or Neutralino mass

Studying Susy-hit to understand more about parameters used in Whizard

### Whizard Simulation(3)



- Γ (stau1) is stau-mass dependent
- Neutralino mass matrix is from GMSB

### Summary

- $(g_{\mu}-2)$  in the MSSM is one of among BSMs that could contain possible candidate for dark matter.
- Investigating how to detect stau pair creation at the ILC.
- Calculating the relation of cross-section with center mass energy, mixing angle and stau-mass (physsim).
- Calculating the relation of stau-mass and cross-section using Whizard.

### Future Plan

- Studying how to generate the signal and reduce BG using Whizard and Marlin.
- More precise analysis of branching ratio and momentum distribution of sleptons for SUSY particles
- Studying of other decay modes of stau.
- Studying of SUGRA and extra dimension (the valuable tests for String theory – AdS/CFT)

### Thank you for listening!

### Back Up

In the MSSM, 2HDM:

At the minimum of the potential

$$H_{u} = \begin{pmatrix} H^{+} \\ H_{u}^{o} \end{pmatrix} \quad H_{d} = \begin{pmatrix} H_{d}^{0} \\ H^{-} \end{pmatrix}$$
$$H_{u} = \begin{pmatrix} 0 \\ v_{u} \end{pmatrix} \quad H_{d} = \begin{pmatrix} v_{d} \\ 0 \end{pmatrix}$$
$$V_{SUSY}^{2} = v_{u}^{2} + v_{d}^{2} \quad \tan\beta = \frac{v_{u}}{v_{d}}$$

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- Neutralino is the combination of  $\widehat{H}_{u}^{o}, \widehat{H}_{d}^{o}, \widehat{B}, \widehat{W}^{o}$  (L-type spinor field) The soft SUSY-breaking mass terms:  $-\frac{1}{2}M_{1}\widehat{B}.\widehat{B}-\frac{1}{2}M_{2}\widehat{W}^{0}.\widehat{W}^{0}+h.c$ ٠

$$\widetilde{G}^{0} = \begin{pmatrix} \widetilde{B} \\ \widetilde{W}^{o} \\ \widetilde{H}_{d}^{o} \\ \widetilde{H}_{u}^{o} \end{pmatrix} \qquad M_{\widetilde{G}^{0}} = \begin{pmatrix} M_{1} & 0 & -\cos\beta\sin\theta_{W}m_{Z} & \sin\beta\sin\theta_{W}m_{Z} \\ 0 & M_{2} & \cos\beta\cos\theta_{W}m_{Z} & -\sin\beta\cos\theta_{W}m_{Z} \\ -\cos\beta\sin\theta_{W}m_{Z} & \cos\beta\cos\theta_{W}m_{Z} & 0 & -\mu \\ \sin\beta\sin\theta_{W}m_{Z} & -\sin\beta\cos\theta_{W}m_{Z} & -\mu & 0 \\ & -23- & & & & & \\ \end{pmatrix}$$

### Stau Reconstruction(1)

• <u>The (mass)<sup>2</sup> matrix for measuring the stau masses</u>

$$\mathbf{M}_{\tilde{\tau}}^{2} = \begin{pmatrix} m_{\tilde{\tau}_{L}}^{2} + m_{\tau}^{2} + \Delta_{\tilde{e}_{L}} & m_{\tau}(A_{o} - \mu tan\beta) \\ m_{\tau}(A_{o} - \mu tan\beta) & m_{\tilde{\tau}_{R}}^{2} + m_{\tau}^{2} + \Delta_{\tilde{e}_{R}} \end{pmatrix}$$

Where,

$$\Delta_{\tilde{e}_L} = \left(-\frac{1}{2} + \sin^2\theta_W\right) m_Z^2 \cos 2\beta$$
$$\Delta_{\tilde{e}_R} = \frac{1}{3} \sin^2\theta_W m_Z^2 \cos 2\beta$$

The M<sup>2</sup> can be diagonalized by the orthogonal transformation

$$\begin{pmatrix} \tilde{\tau}_1 \\ \tilde{\tau}_2 \end{pmatrix} = \begin{pmatrix} \cos\theta_{\tau} & -\sin\theta_{\tau} \\ \sin\theta_{\tau} & \cos\theta_{\tau} \end{pmatrix} \begin{pmatrix} \tilde{\tau}_L \\ \tilde{\tau}_R \end{pmatrix}$$
  
The eigenvalues are:  $m_{\tilde{\tau}_1}^2, m_{\tilde{\tau}_2}^2 (m_{\tilde{\tau}_1}^2 < m_{\tilde{\tau}_2}^2)$   
This is what we are looking

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for!

### Stau Reconstruction(2)

The cross-section for stau pair production: