

# AXEL

-A Xenon Electro Luminescence detector -  
High Density Xenon gas TPC  
for double beta decay search and direction  
sensitive dark matter search

- References are from “Noble Gas Detectors” unless otherwise noted.

# Double Beta decay candidates

|       | abound(%) | $\tau(2\nu\beta\beta)$ yr | Q(keV) | $Q^5/1E16$ | $Q^5 \times \tau(2\nu\beta\beta)/1E36$ |                      |
|-------|-----------|---------------------------|--------|------------|--|----------------------|
| 48Ca  | 0.187     | 3.9E+19                   | 4271   | 142.12     | 55.4                                   | enrichment difficult |
| 76Ge  | 7.8       | 1.7E+21                   | 2039   | 3.52       | 59.9                                   |                      |
| 82Se  | 9.2       | 9.6E+19                   | 2995   | 24.10      | <b>23.1</b>                            |                      |
| 96Zr  | 2.8       | 2.0E+19                   | 3350   | 42.19      | 8.4                                    |                      |
| 100Mo | 9.6       | 7.1E+18                   | 3034   | 25.71      | <b>1.8</b>                             |                      |
| 110Pd | 11.8      |                           | 2013   | 3.31       |  |                      |
| 116Cd | 7.5       | 2.8E+19                   | 2802   | 17.27      | 4.8                                    |                      |
| 124Sn | 5.64      |                           | 2228   | 5.49       |  |                      |
| 130Te | 34.5      | 7.6E+20                   | 2529   | 10.35      | 78.6                                   |                      |
| 136Xe | 8.9       | 2.2E+21                   | 2479   | 9.36       | <b>208.8</b>                           |                      |
| 150Nd | 5.6       | 9.2E+18                   | 3367   | 43.27      | 4.0                                    | enrichment difficult |

\*  $\tau(2\nu\beta\beta) \propto Q^{11}$ ,  $\tau(0\nu\beta\beta) \propto Q^5$

# $^{136}\text{Xe}$

- Xenon production rate
  - 5000-7000m<sup>3</sup>/yr ~50ton in 1998.
- 2x10<sup>6</sup> kton exists in air
- Enrichment is relatively easy.

|    | Density at 0°C,1atm (g/L) | Liquid Density (g/cm <sup>3</sup> ) | Melting point(K) | Boiling point(K) |
|----|---------------------------|-------------------------------------|------------------|------------------|
| He | 0.1786                    | 0.145                               | -                | 4.22             |
| Ne | 0.9002                    | 1.2                                 | 24.56            | 27.07            |
| Ar | 1.784                     | 1.398                               | 88.80            | 87.30            |
| Kr | 3.749                     | 2.413                               | 115.79           | 119.93           |
| Xe | 5.894                     | 3.053                               | 161.4            | 165.03           |

# Energy resolution

- Statistical limit
  - W-value 21.5 eV, Fano factor < 0.17
    - 0.29%(FWHM)@2.48MeV (0.55%(FWHM)@662keV)
- At higher density, energy resolution becomes worse. → reject liquid option.

*A. Bolotnikov, B. Ramsey / Nucl. Instr. and Meth. in Phys. Res. A 396 (1997) 360–370*

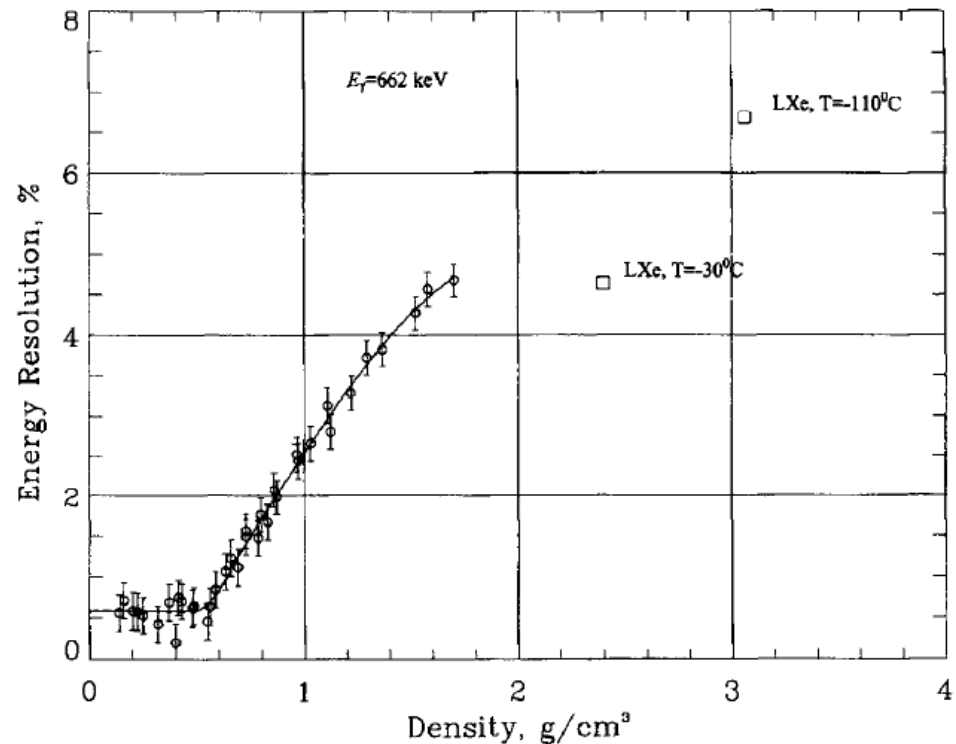


Fig. 5. Density dependencies of the intrinsic energy resolution (%FWHM) measured for 662 keV gamma-rays.

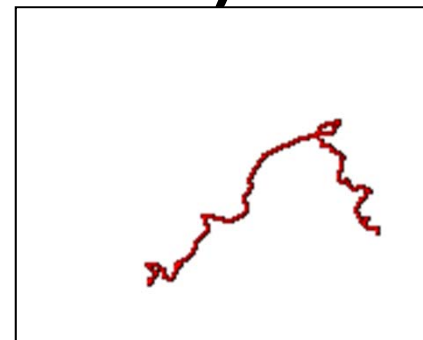
# High density (15~30 x STP)

## High pressure

- Need pressure vessel including feed through
- Need photo detector operatable at high pressure

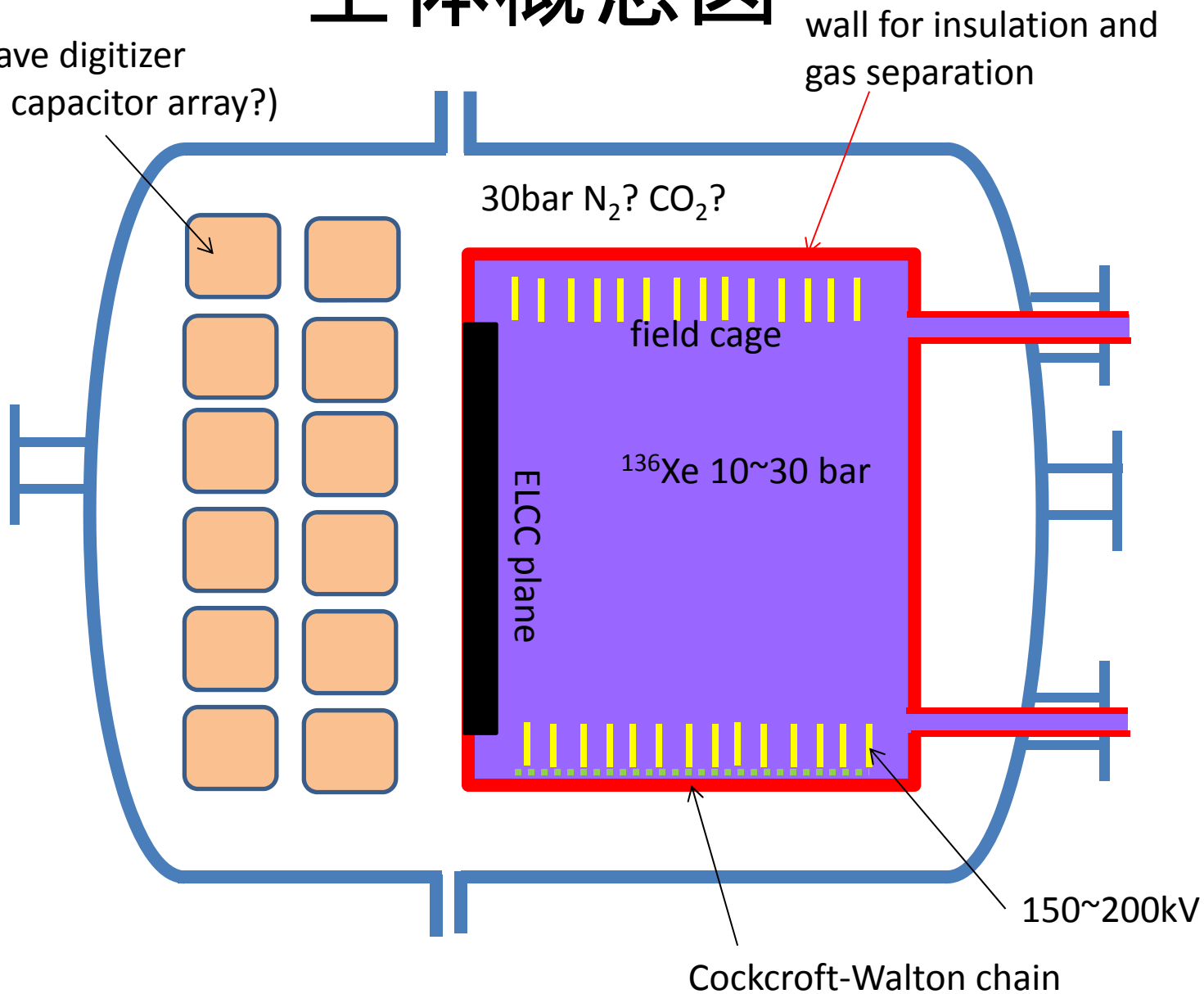
# 我々のBaseline design(目標)

- 1 ton enriched  $^{136}\text{Xe}$  gas (not liquid)
- At 15~30 times higher density than STP
  - $\rho = 0.088\sim 0.18\text{g/cm}^3$
  - e.g.  $\phi 2\text{m}\times 1.7\text{m(H)}$  cylinder at  $0.18\text{g/cm}^3$
- Use proportional scintillation mode (Electroluminescence) for energy measurement
  - Energy resolution goal  $< 0.5\%$ (FWHM)
  - Ultraviolet photon( $\sim 170\text{nm}$ ) detection by MPPC
- Tracking as TPC
  - Range( $2.5\text{MeV e}$ )  $\sim 210\text{ cm}$  at STP
  - $T_0$  by primary scintillation signal
  - Sample 15~20 points using pads.  $5\sim 7.5\text{mm}$  spacing  $\rightarrow 5.5\times 10^4\sim 1.2\times 10^5\text{ ch}$
  - Purpose is to identify two blobs at track ends.  $\rightarrow$  distinguish from  $\alpha$ 's and  $\gamma$ 's.
  - Electric field for drift :  $\sim 1.5\text{kV/cm}$ @30bar  $\rightarrow$  drift velocity  $\sim 1\text{m/ms}$
- Energy measurement by ELCC(see next pages)



# 全体概念図

~5MHz wave digitizer  
(Switched capacitor array?)





# Milestones

- MPPC 64チャンネル、10気圧検出器を製作し  
1.3MeVガンマ線の測定で原理証明
  - このための開発要素、盛りだくさん。
- 来年度以降の目標は
  - キセノン10気圧 9kg検出器
    - ↓
  - 30気圧 27kg検出器で世界記録更新
    - ↓
  - 30気圧1トン検出器で発見！

# Scintillation

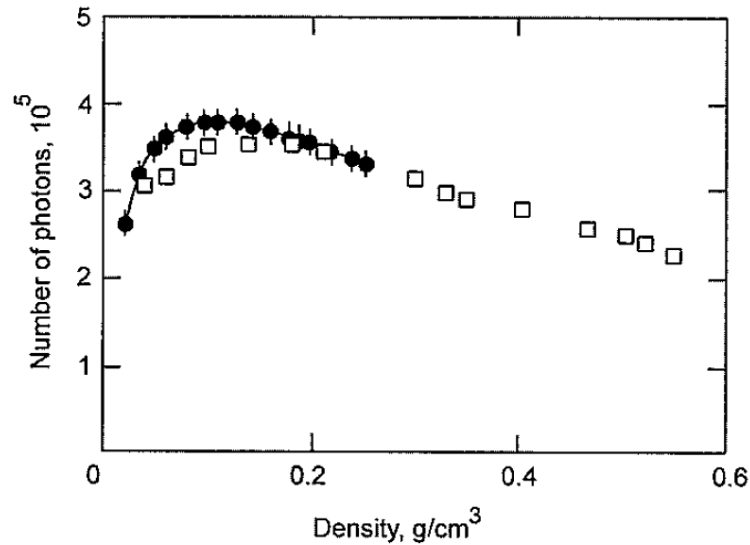


Fig. 3.38 Dependence of light output of scintillations excited by alpha particles on the density of high-pressure xenon: open squares represent data of Bolotnikov and Ramsey [33], closed circles represent data of Kobayashi et al. [198]. Redrawn from [198].

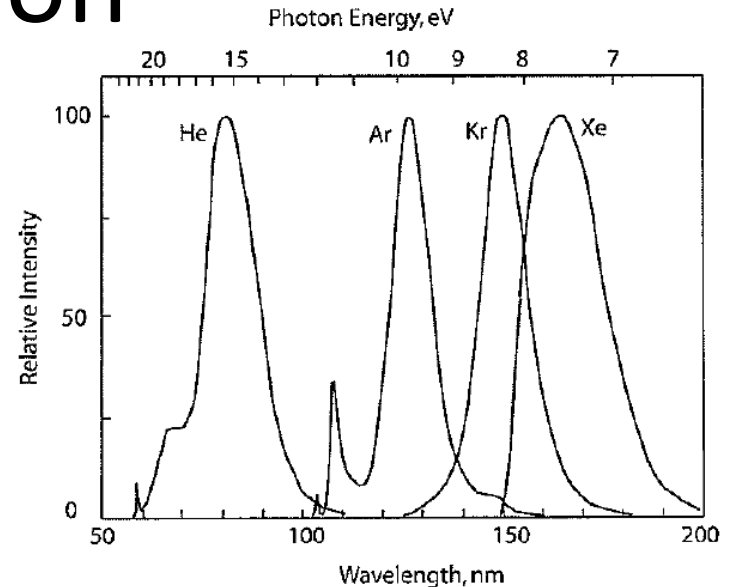
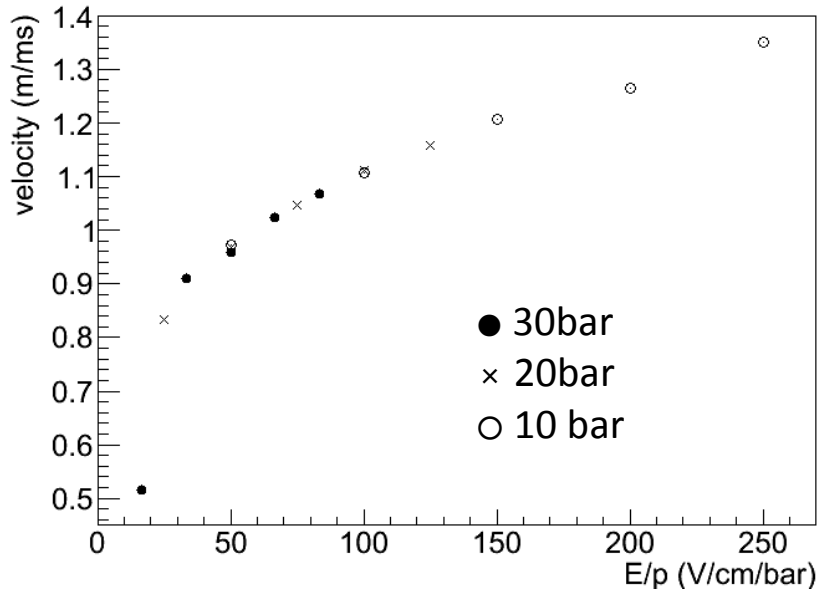


Fig. 3.27 Noble gas continua of helium, argon, krypton, and xenon normalized to the same intensity of the maximum. Redrawn from [176].

6. *Impurity emission.* Any impurity molecules present may quench the emission of hard UV light, and their own emission may be stimulated by transfer of excitation energy. This is of practical importance in the xenon, which is used as a wavelength-shifter the emission from helium or argon, which excitation energies are higher than the minimum xenon excitation level. At relative concentration of  $10^{-5}$ , nitrogen is adequate for efficient energy transfer from all the noble gases at atmospheric pressure. At relative concentration of  $10^{-3}$ , Xe can be used as a wavelength shifter, for example in  $^3\text{He}$  scintillator at 3.5 MPa pressure [171]. Sometimes, impurities have nonradiative transitions or emit in regions where the sensitivity of photodetectors is limited. This effect may dramatically reduce the observed light yield of noble gas scintillators.

What is the influence on scintillation from  $\text{N}_2$  or  $\text{H}_2$  addition?

# Electric Field and electron drift



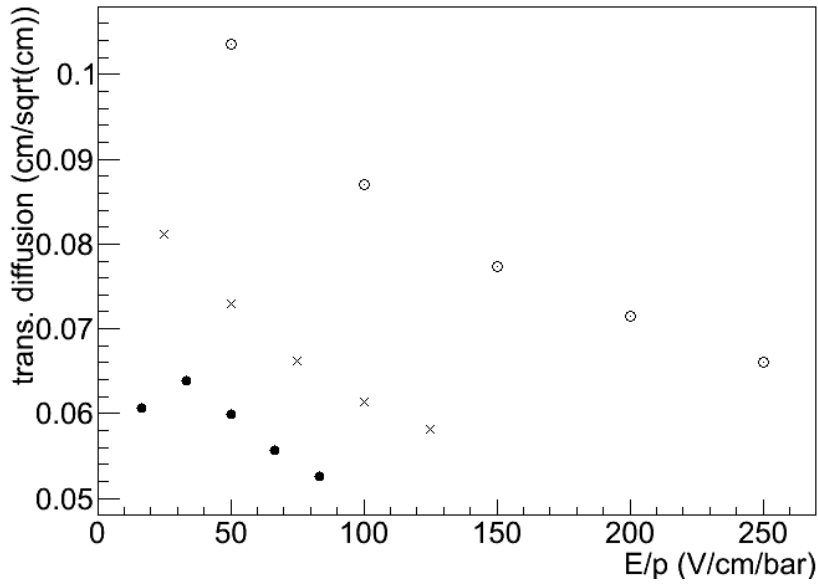
@  $E=1.5\text{kV/cm}$ ,  $p=30\text{ bar}$

Drift velocity  $0.96\text{m/ms}$

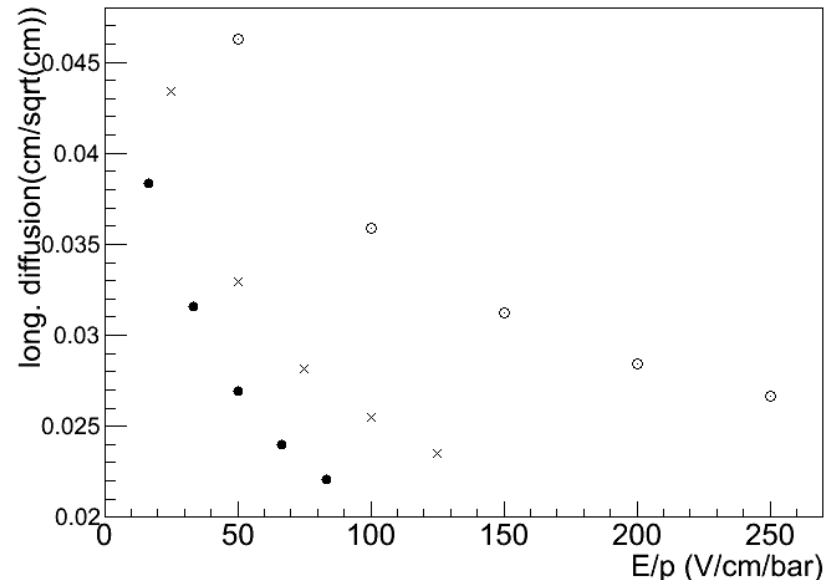
- Possibly add  $\text{H}_2$ ,  $\text{N}_2$  or He to increase the drift velocity.
  - It will also reduce diffusion.
  - It will also reduce light yield

diffusion after 1m drift

- transverse : 6mm
- longitudinal : 2.7mm

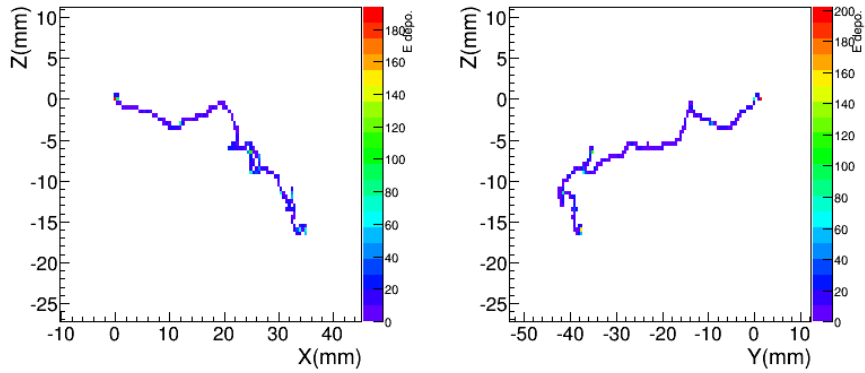


MAGBOLTZ calculation

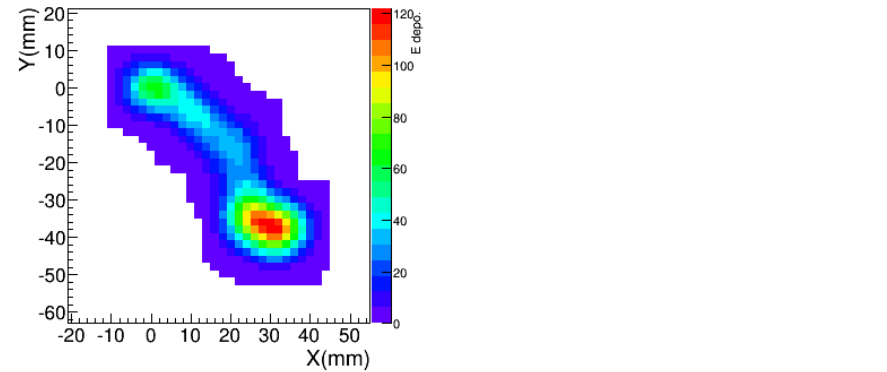
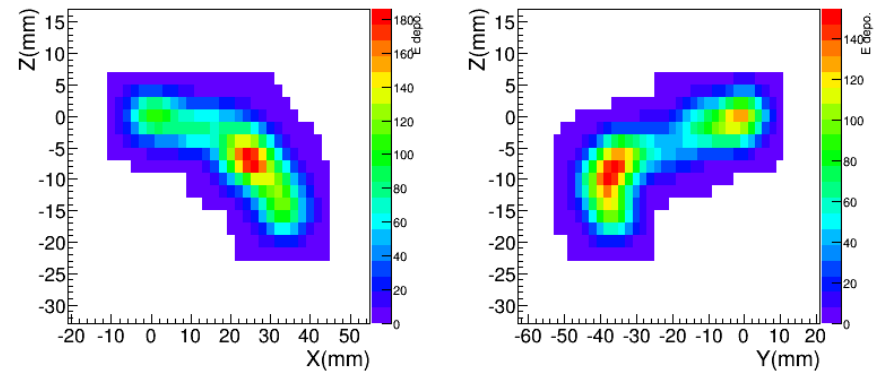
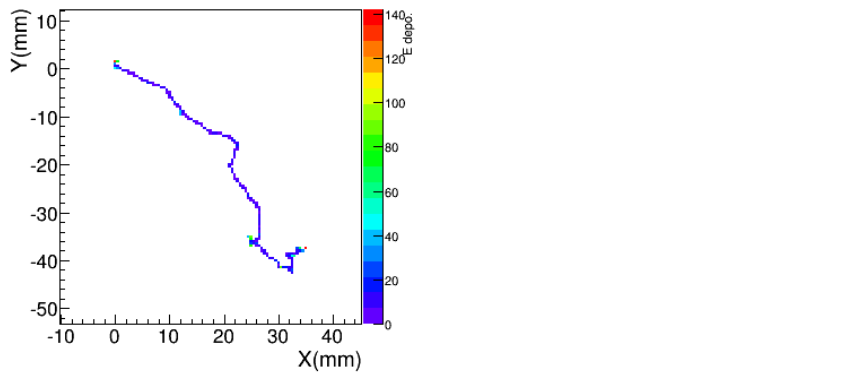


# simulation example 1

initial ionization distribution



after 1m drift



# Electroluminescence

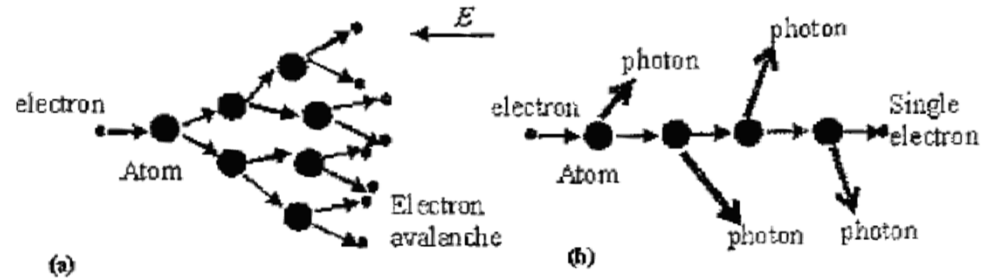


Fig. 6.2 Amplification process in gas detectors with gas gain (a) and electroluminescence (b) or proportional scintillation.

Good and stable linearity are expected because

- A linear amplification process.
  - c.f. Electron multiplication process is exponential process.
- The number of produced photons is proportional to the voltage drop rather than to the field strength.
  - Insensitive to the capacitance change by microphonic vibration

To keep original resolution determined by career generation,

- $\epsilon Y > 1/F$  ( $\epsilon > 5\%$  for 400 photons/e)

c.f. usually  $< 1\%$

In a uniform electric field, the number of photons generated by one drifting electron  $N_{ph}$ , is proportional to the drift path  $x$  [cm] and for xenon gas at room temperature may be defined via the reduced electric field strength  $E/p$  [kV cm<sup>-1</sup> bar<sup>-1</sup>] and the gas pressure  $p$  [bar] using the empirical equation [144] as

$$dN_{ph}/dx = 70(E/p - 1.0)p \quad (\text{UV photons/e cm drift}) \quad (3.24)$$

The intensity of electroluminescence (EL) up to 1700 photons/cm in xenon at 0.5 MPa pressure has been demonstrated [144]. Taking into account that the energy of a single photon of 172 nm wavelength is about 8.4 eV, one can calculate the efficiency of conversion of the energy of the electric field into the photon emission:  $\xi = (1700 \times 8.4) / (3400 \times 5) = 84\%$ . The rest of the energy ac-

w/ 5 mm gap  
 To get 200 (400) photons,  
 18 (21) kV@30atm.  
 7.9 (11) kV@10atm.  
 3.4 (6.2) kV@1atm.

w/ 4 mm gap  
 To get 200 (400) photons,  
 4.9 (7.7) kV@5atm.  
 3.3(6.1) kV@1atm.

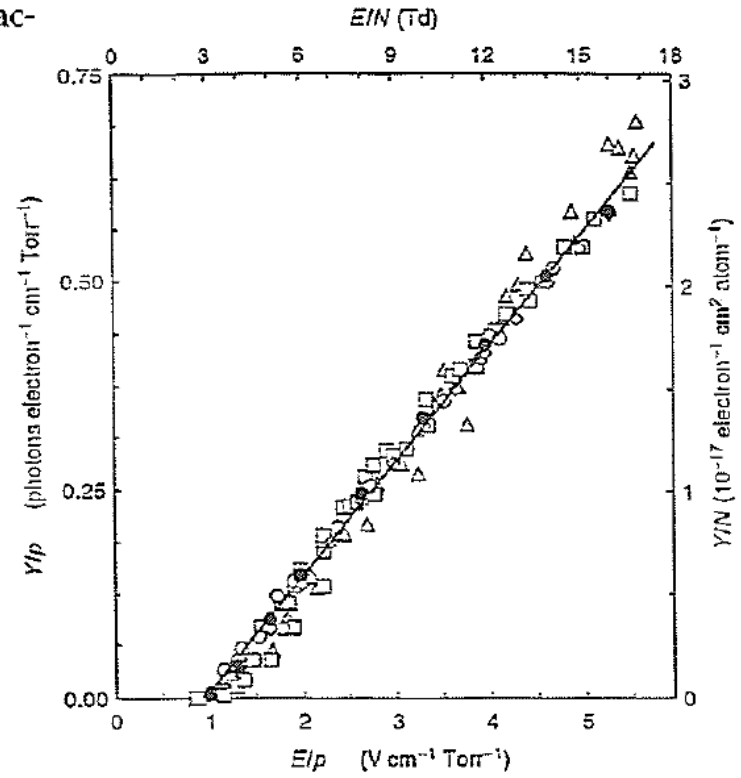
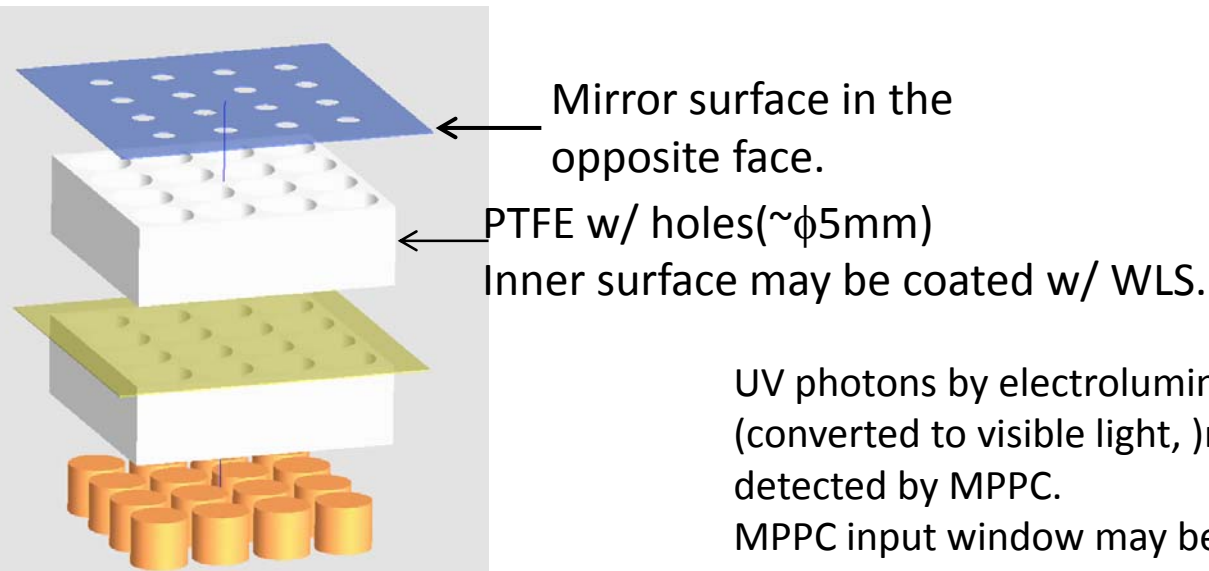
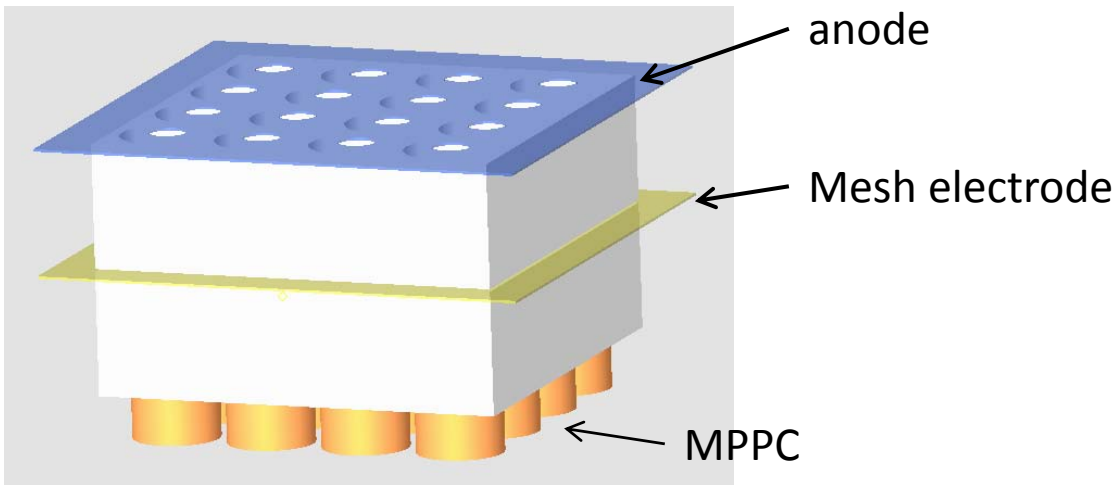
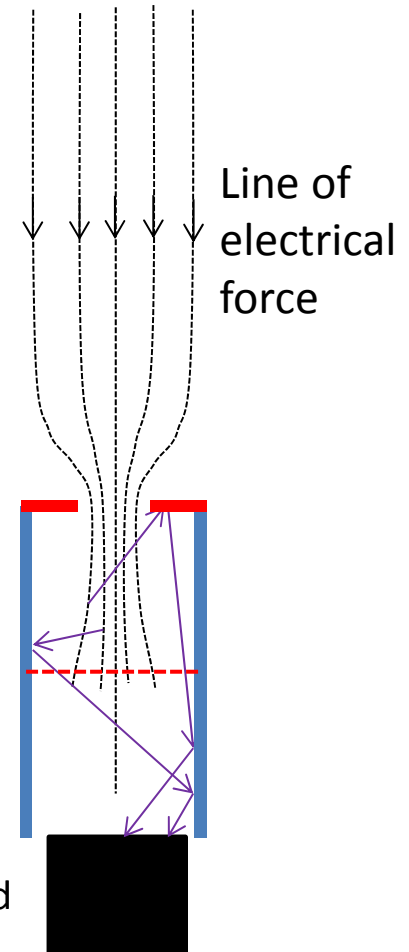


Fig. 3.17 Reduced light output of electroluminescence of xenon gas at 293 K temperature and normal pressure as a function of the reduced electric field strength (compilation of experimental and computer simulation data by Conde [143]).

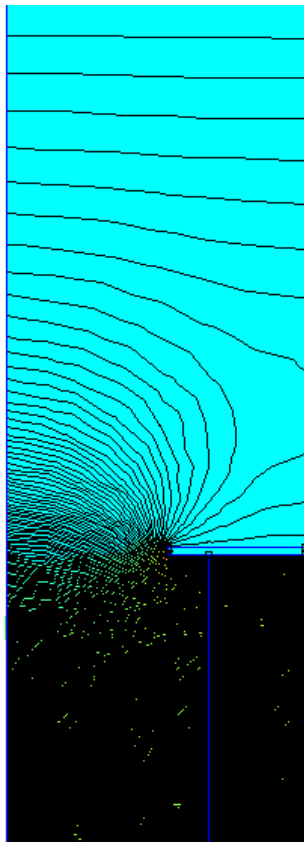
# Readout by Electroluminescence light collection cell (ELCC)



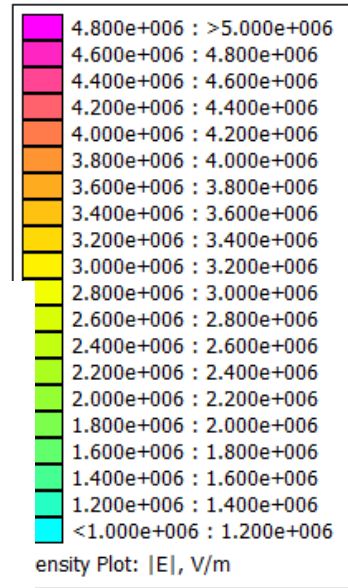
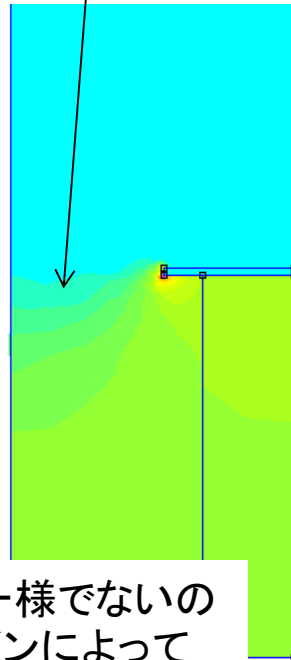
UV photons by electroluminescence are (converted to visible light, ) reflected and detected by MPPC.  
MPPC input window may be coated w/ WLS.



# Electric Field in ELCC



EL threshold field strength

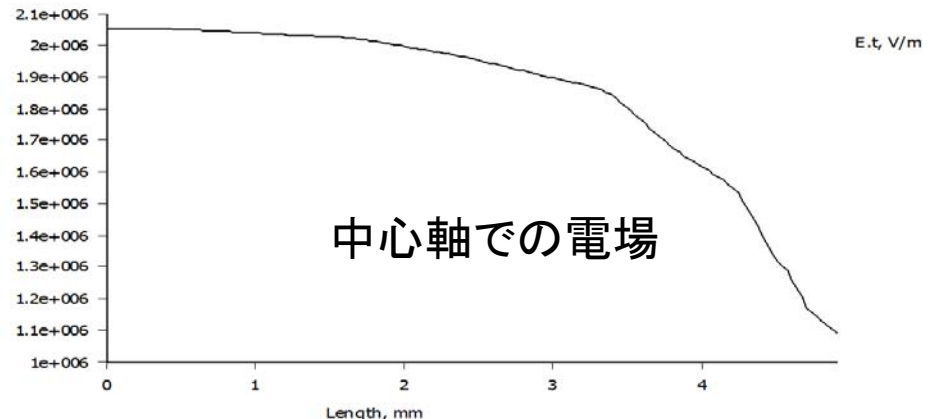


10気圧の場合。  
アノードの穴はφ4。  
PTFEの穴はφ5。  
anodeは10.5kV  
ドリフト電場は  
500V/cm

電場の強さが場所によって一様でないの  
で、EL光が電子のドリフトラインによっ  
てばらつく可能性がある。

詳細なシミュレーションをしたいが、どうい  
うツールがある？

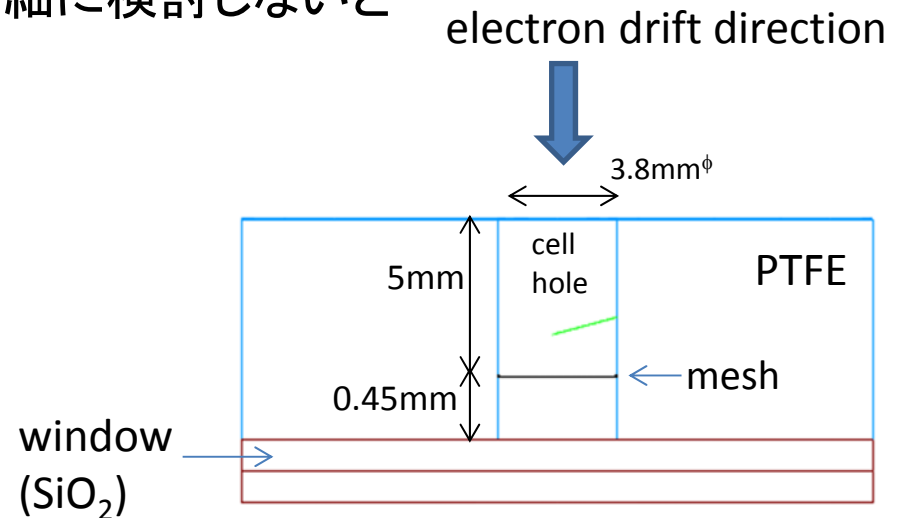
(FEMMでscriptを書いて電場マップを作り、  
geant4?)





# ELCC light collection efficiency

- by Geant4. cell内で光を一様発生
- PTFEについては、波長175nm以下での反射が観測されていない。175nm以上では反射率は>55%とされているので、平均反射率として11%を仮定。
- 収集効率は12.9%（これに、光検出器のefficiencyを掛ける必要あり。）
- ギャップを長くすれば、EL gainは十分に稼げるが、電圧は高くなる。wavelength shifterを塗って可視光にしてしまう手もある。どちらが良いかは、詳細に検討しないとわからない。



# MPPCに対する要求

## EL光

- 全光量は $\sim 1e6$  photons.(ELで10倍増幅した場合)
- MPPC1個での瞬間最大光量は $1e3$  photons/50ns
- 3mm角、ピクセルピッチ $10\mu\text{m}$ (ピクセル数 $9e4$ )からピクセルピッチ $25\mu\text{m}$ (ピクセル数 $1.4e4$ )のMPPC
- 保護膜なしのタイプで紫外光を直接検出するか、WLSで可視光に変換するか。

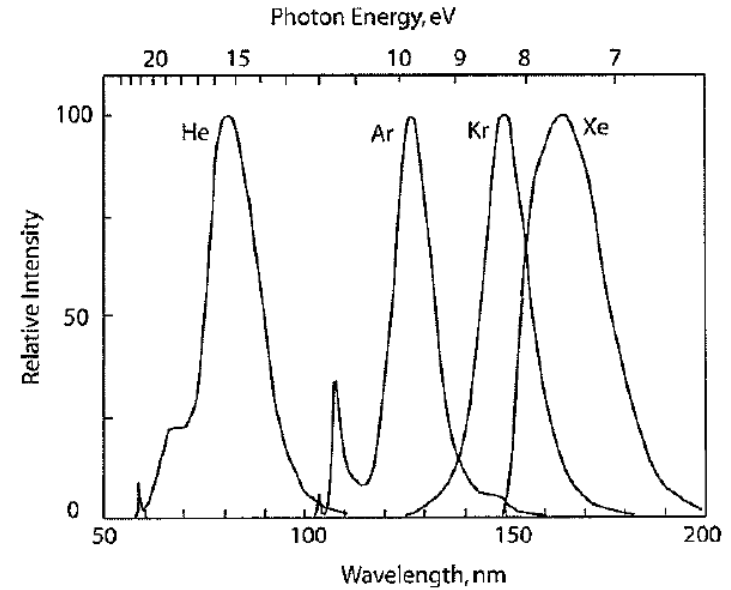


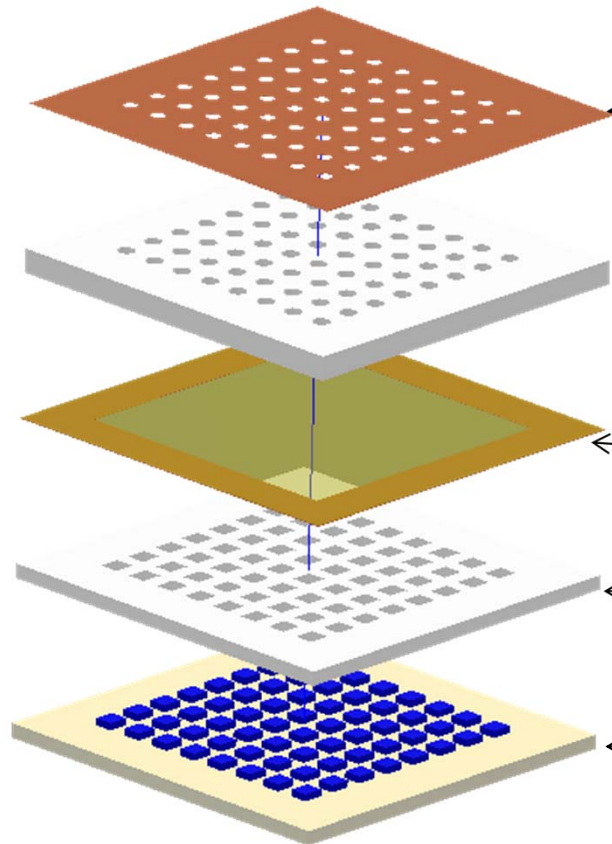
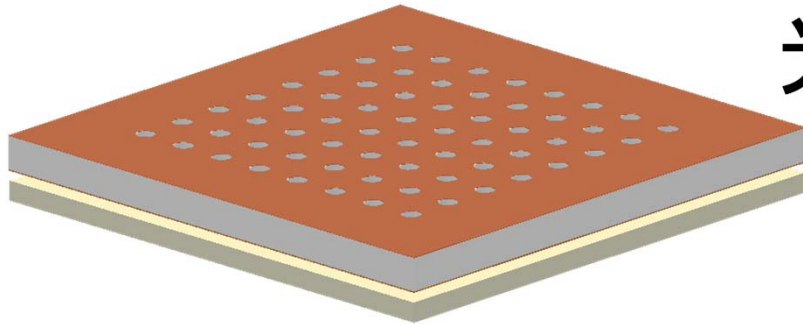
Fig. 3.27 Noble gas continua of helium, argon, krypton, and xenon normalized to the same intensity of the maximum. Redrawn from [176].

“Noble Gas Detectors”より

[176] R.E.Huffman, J.C.Larrabee and Y.Tanaka, Appl. Opt. 4, 1581-1588 (1965)

# 光収集セル(ELCC)の構造

- 電極1と電極2の間に高電圧がかかる。  
電極2はGND



電極1 無酸素銅  
厚さ~100 $\mu$ m

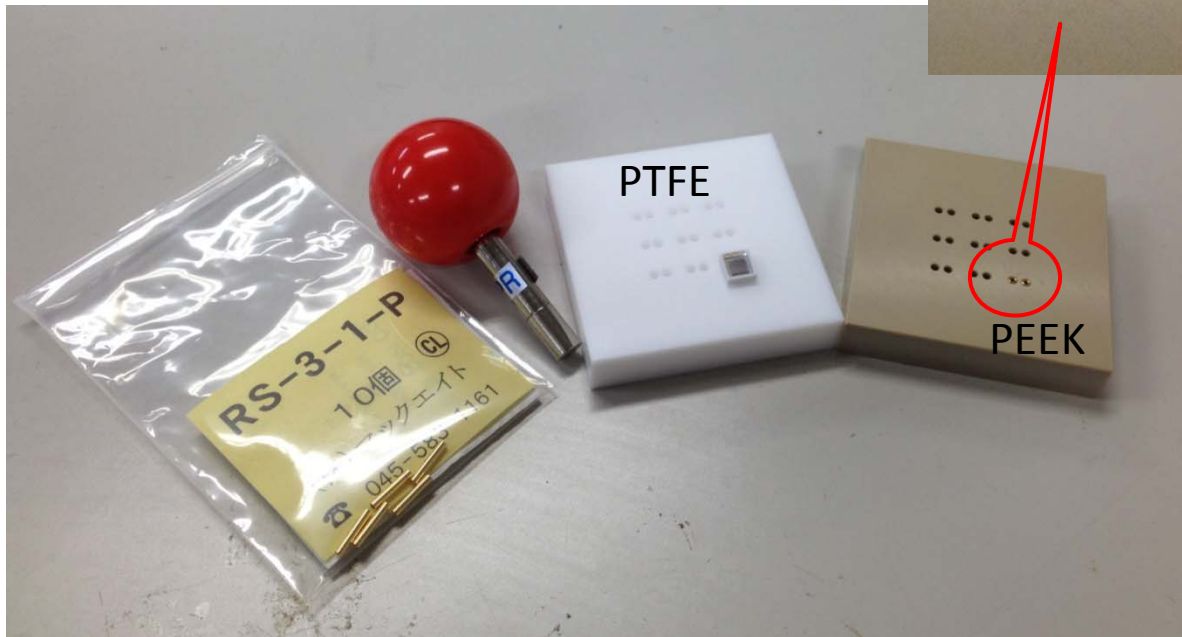
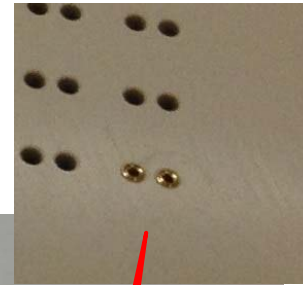
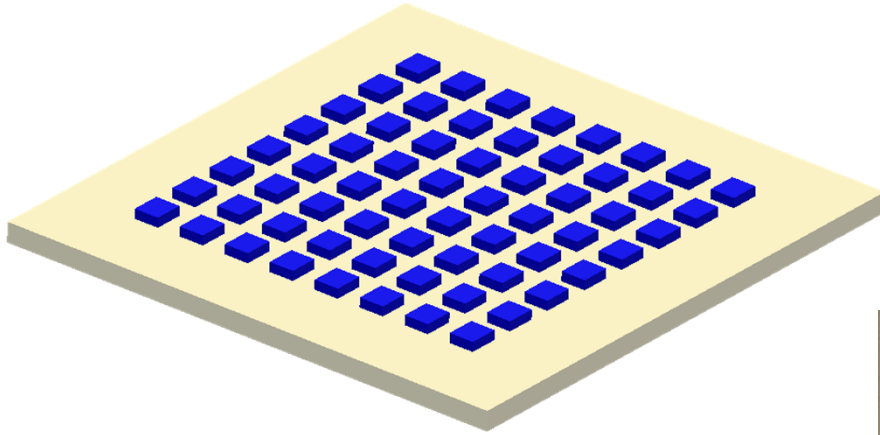
PTFE(テフロン) (紫外光の反射材の役割)  
厚さ5mm

電極2 メッシュ

PTFE(テフロン) (紫外光の反射材の役割)  
厚さ2mm?  
MPPCが入り込むような穴が開いている。

MPPC plane  
3mm角MPPCを7.5mmピッチで並べる

# MPPC plane



プラスチックの板に両方向  
中継ソケットを差し込む。  
ソケットの表側にMPPC、片  
側にケーブルを圧着したピ  
ンを差し込む。  
PTFEだと、MPPCを抜くとき  
に、ソケットが板から抜けて  
しまった。

# primary scintillation to determine $t_0$

- 全光量
  - $W_s = 76\text{eV}$  (from NEXT CDR)  $\rightarrow$  32,300 photons
- Tracking planeでのlight yield
  - $2\text{m}\phi \times 1.7\text{mH}$ の円柱で両側にtracking planeがあると仮定  $\rightarrow$  planeの acceptance  $17.5\% \times 2$
  - ELCC開口率
    - $2\text{mm}\phi$ ,  $5\text{mm}$  spacingだと12.6%
  - PDE 20%と仮定
  - $32,300 \times 0.175 \times 2 \times 0.126 \times 0.2 = 284 \text{ p.e.}$
  - ELCCの開口穴半径を $3\text{mm}\phi$ にすると、639p.e.
- これを全MPPC( $>1\text{e}5$  ch)の和として検出するのは難しい。(dark currentと区別がつかない)
- 低温にしてMPPCのdark rateを下げるか高圧化で動作するPMTが必要。

# Wavelength shifter

**Tab. 4.1** High-pressure noble gas scintillation detectors.

|                 | Pressure (MPa),<br>size(cm) | WLS               | Particle,<br>Energy(MeV) | En. Res.,<br>% FWHM | Ref.  |
|-----------------|-----------------------------|-------------------|--------------------------|---------------------|-------|
| <sup>4</sup> He | 13.8; 5 I.D.×8.6            | 7%Xe, DPS         | n, 1.0                   | 36                  | [234] |
| Ne              | 1.7                         | 20%Xe             | n, 3.5                   | 6.7                 | [231] |
| <sup>3</sup> He | 2.0                         | 10%Xe             | n, < 3                   | 14                  | [232] |
|                 | 3.5; 1.7×0.5                | 0.5%Xe, p-TP      | n <sub>th</sub>          | 18                  | [171] |
|                 | 24.8; 38 cm <sup>3</sup>    | DPS               | n <sub>th</sub>          | 54                  | [233] |
|                 |                             |                   | α, 5.15                  | 16                  |       |
|                 | 20.0; 4 I.D.×4              | 5%Xe, p-TP        | n, 2.5                   | 4.8                 | [236] |
|                 | 13.8; 5 I.D.×8.6            | 2%Xe, DPS         | n <sub>th</sub>          | 31                  | [235] |
| Xe              | 2.8                         | 25%N <sub>2</sub> | n, 0.5                   | 29.5                | [231] |

Note: WLS - wavelength shifting; DPS - trans p,p'-diphenylstilbene; p-TP - paraterphenyl; n<sub>th</sub> - thermal neutrons.

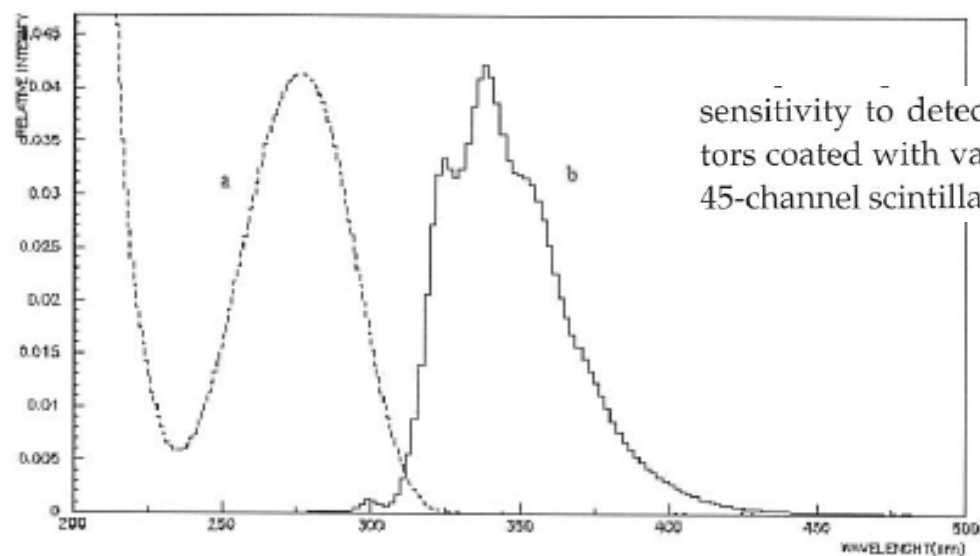
- Acrylic tube coated with a thin layer of polysterene doped with organic fluor TPB ( tetraphenyl butadiene) (p.114)
- N<sub>2</sub> output peak 340nm

### 8.5.2.1 Wavelength Shifters Dissolved in Noble Gases

One of the first wavelength shifters used with noble gas scintillators was a gas admixture of nitrogen. For example, Grün and Schopper [228] used argon in mixture with 2% of nitrogen in their development of a detector for triggering of cloud chambers. The addition of a small amount of nitrogen to noble gases enhances the light emission in the blue range. At earlier studies it was concluded that nitrogen acts as a simple fluorescent converter. However, the nitrogen also acts as the quenching agent: its addition to the noble gas results in decreasing the absolute scintillation efficiency. For example, Northrop and Nobles [13] observed that there is a reduction in the practical light output of a xenon gas scintillator, used with solid wavelength shifter, when nitrogen is added. The efficiency is reduced by about 1/3 by addition of 10% of nitrogen or hydrogen. The latter is often used in high-pressure xenon detectors to increase the drift velocity of electrons. Another undesirable feature of nitrogen is the introduction of a slower component of decay time. For example, the addition of  $10^{-4}$  N<sub>2</sub> introduces in argon scintillator a component with decay time of 0.5  $\mu$ s, which accounts for 75% of the total photon emission [170]; the decay time reduces with an increase in the concentration of nitrogen, however, this reduces the total light yield. From all these facts, one can conclude

### 8.5.2.2 Solid Wavelength Shifters

A number of solid wavelength shifters deposited onto optical elements such as windows and mirrors have been considered including *trans*-stilbene, tetraphenylbutadiene (TPB), sodium salicylate, p-quaterphenyl, diphenylstilbene, and p-terphenyl (see, for example, McKinsey et al. [429] and references therein). The last one was found to be the best one to be used with xenon and xenon-containing gas mixtures from the point of view of high quantum efficiency ( $> 90\%$  according to Belogurov et al. [144]), low hydroscopic, chemical inertness, and exclusive radiation hardness. Emission and absorption spectra of p-terphenyl are presented in Fig. 8.10.



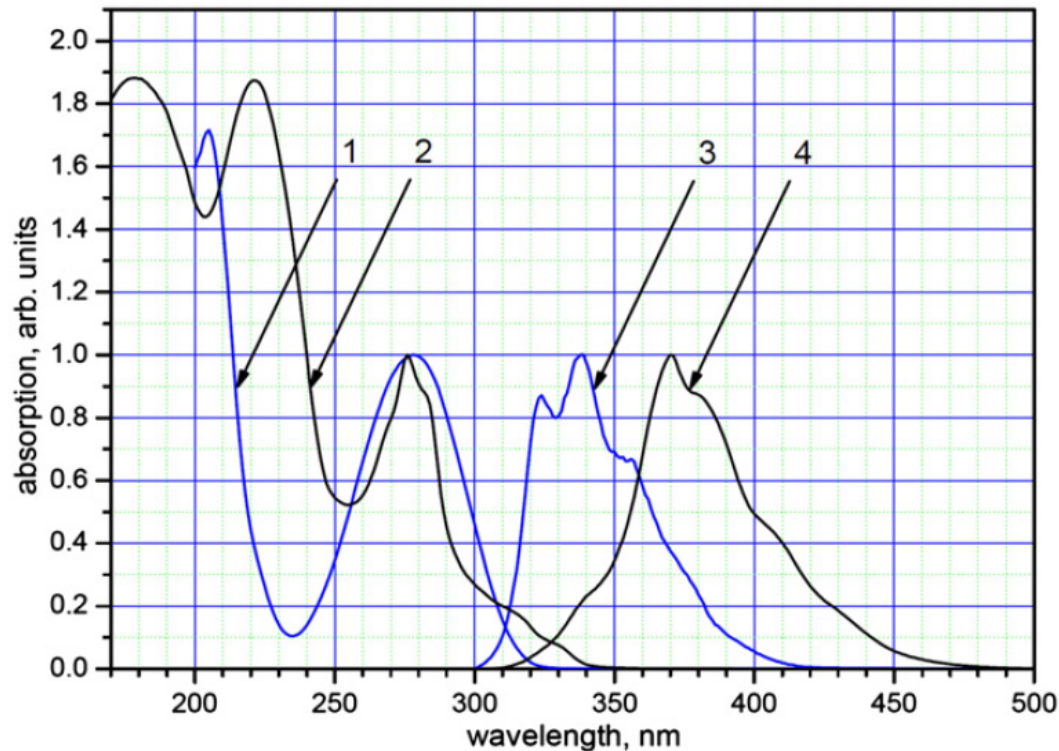
sensitivity to detection of Cherenkov light [431]. Aluminized Mylar reflectors coated with vacuum deposited p-TP have been used in construction of a 45-channel scintillation LXe/LKr electromagnetic calorimeter (see Chapter 4).

**Fig. 8.10** (a) Absorption and (b) emission spectra of p-terphenyl [407].

The p-terphenyl is reported to have a short decay time of 2–5 ns and widely used as a scintillating dye in plastic scintillators. Kumar and Datta [430] com-



# Wavelength shifter – *p*-terphenyl



**Fig. 1.** Absorption and emission spectra of *p*-terphenyl. 1 – *p*-terphenyl in solvent, absorption, 2 – polycrystalline *p*-terphenyl, absorption, 3 – *p*-terphenyl in solvent, emission, 4 – polycrystalline *p*-terphenyl, emission.

D.Yu.Akimov et. al, “Development of VUV wavelength shifter for the use with a visible light photodetector in noble gas filled detectors”

<http://dx.doi.org/10.1016/j.nima.2011.12.036>

D.Yu. Akimov, et al., Nucl. Instr. & Meth. A (2011),  
doi:10.1016/ j.nima.2011.12.036

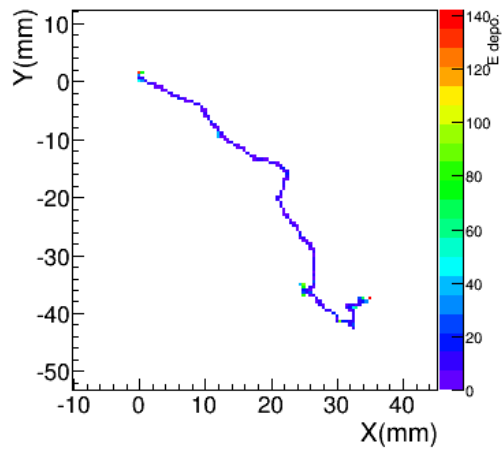
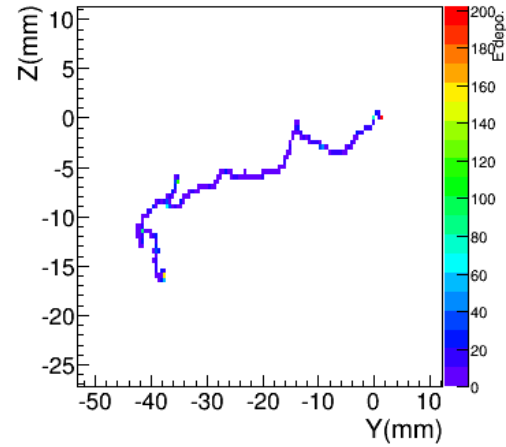
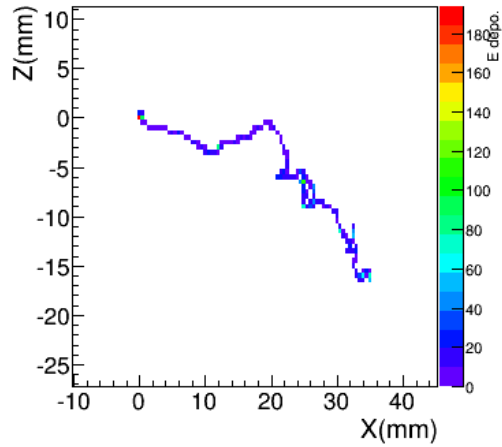
A *p*-terphenyl is known to be a quite volatile organic substance. The use of it is seriously problematic in the liquid noble gas detectors with charge collection, especially in the liquid xenon ones, because it dramatically reduces the lifetime of free electrons created by an ionising particle. To avoid contamination of the liquid xenon by the *p*-terphenyl molecules are coated the WLS with a  $\sim 1 \mu\text{m}$  poly-para-xylylene protection film. Poly-para-xylylene (Parylene N) [16] is chosen due to its well known properties such as the very low permeability to gases, and the possibility to form a conformal optically transparent film practically free of pin-holes even for the thicknesses down to several tens Å. Since the poly-para-xylylene is not transparent for the UV light, the *p*-terphenyl is deposited to the 1- mm thick sapphire, which is situated the first on the way of the UV light. The photodetector then to be situated

# TPB (tetraphenyl butadiene) coating

- <http://microboone-docdb.fnal.gov/cgi-bin/ShowDocument?docid=1006>
- “Environmental Effects on TPB Wavelength-Shifting Coatings”, C.S. Chiu et.al, JINST 7 (2012) P07007
  - a TPB solution consisting of a 1:1:43 ratio by mass of PS to TPB to toluene. Each acrylic plate required three coatings of the TPB solution. This has been measured to have about 50% of the efficiency of the evaporative coating.
  - the harmful effect of certain wavelengths of light on the degradation of 50% TPB-PS chemical TPB coatings and isolated the most damaging wavelength range to the UV spectrum. On the other hand, we have seen that humidity does not play a significant role in long run plate degradation.
- Additional info. From the top document
  - Above 33%, the TPB will crystallize out of solution.
  - Adding ethanol to this mixture breaks up the surface tension, making the coating look smooth and uniform
  - Plates are wiped after coating to remove any loose TPB, after which they do not degrade further and are very resilient

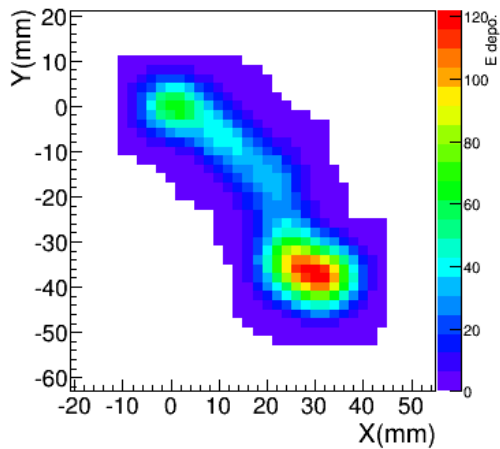
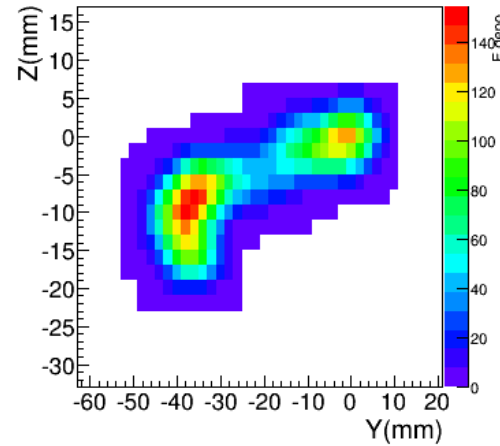
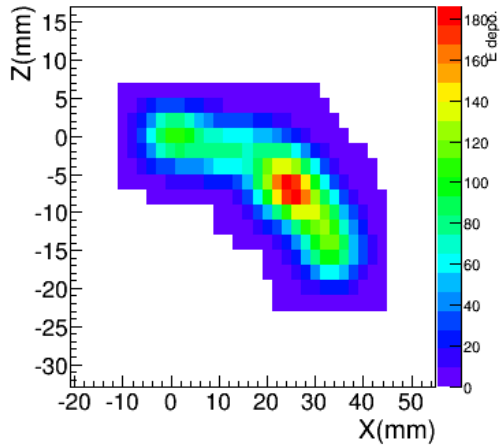
# simulation example 1

## row track



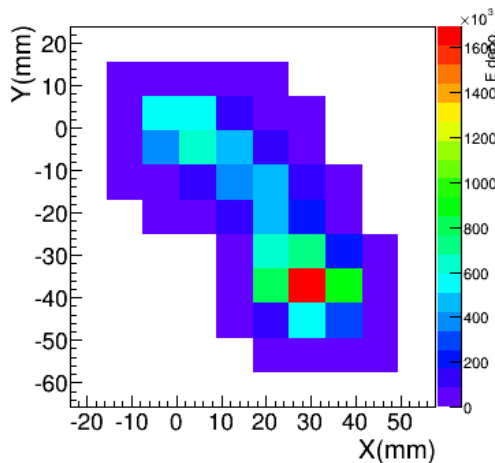
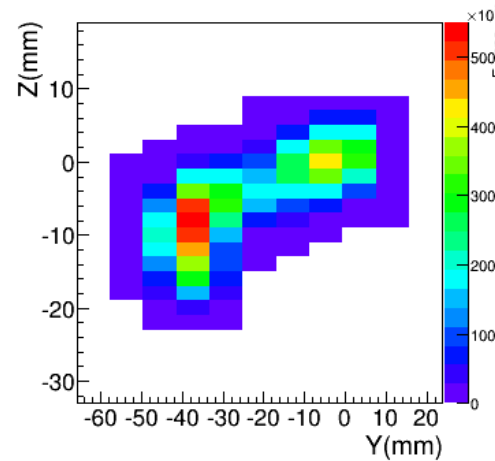
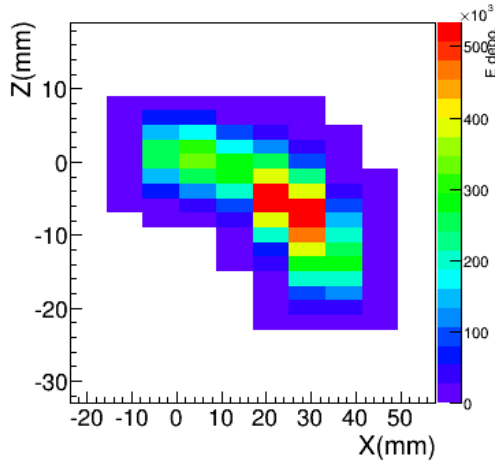
# simulation example 1

## diffusion after 1m drift



# simulation example 1

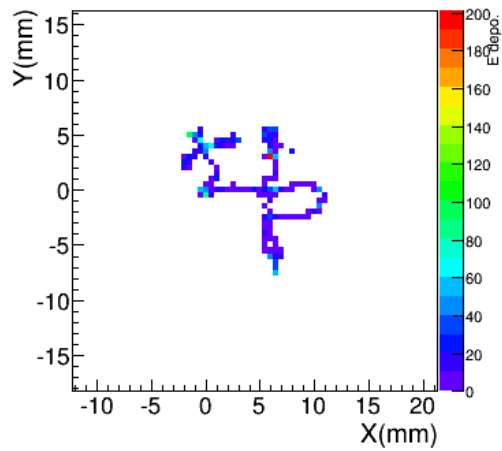
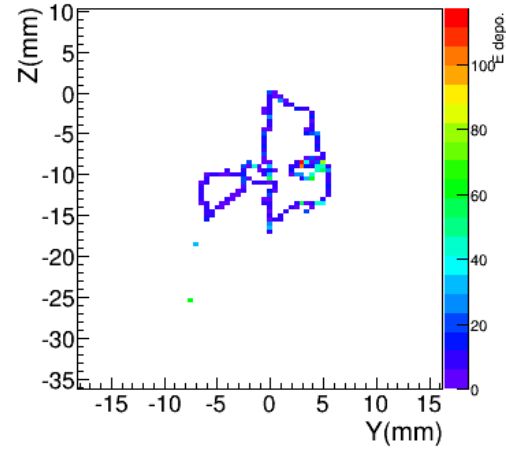
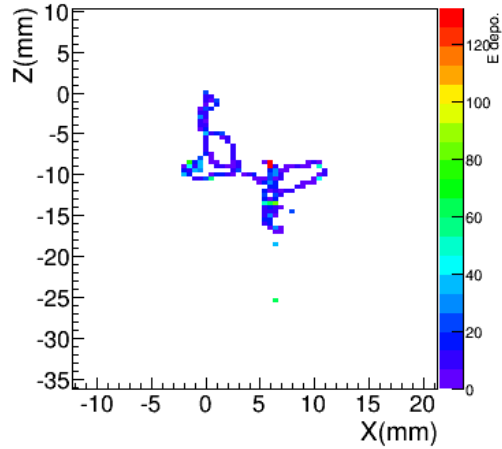
segmentation (7.5mm in x or y and 2mm in z)



z marginalized by  
ELCC generation  
time(4 $\mu$ s)

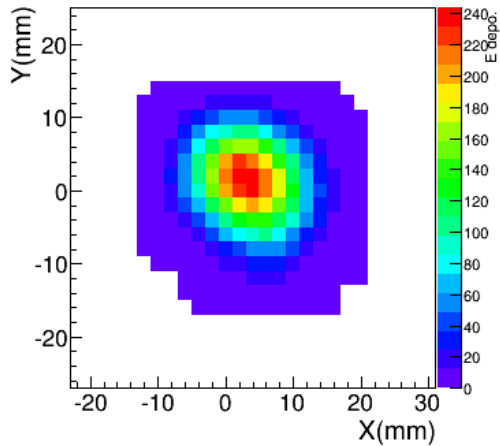
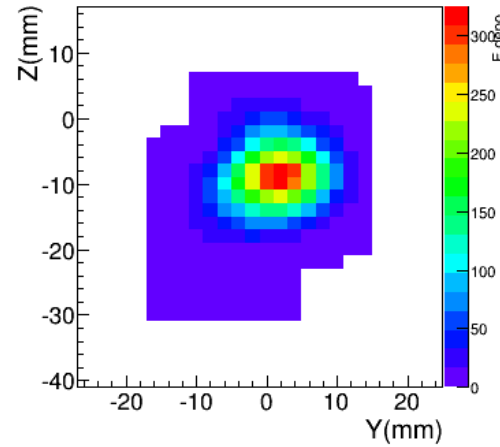
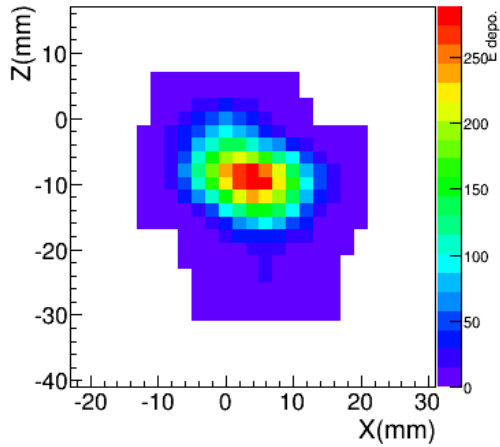
# simulation example 2

## row track



# simulation example 2

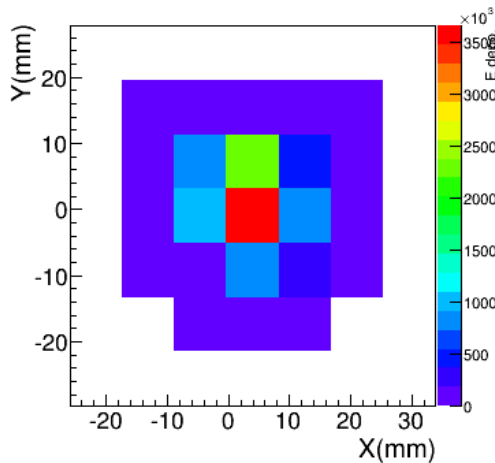
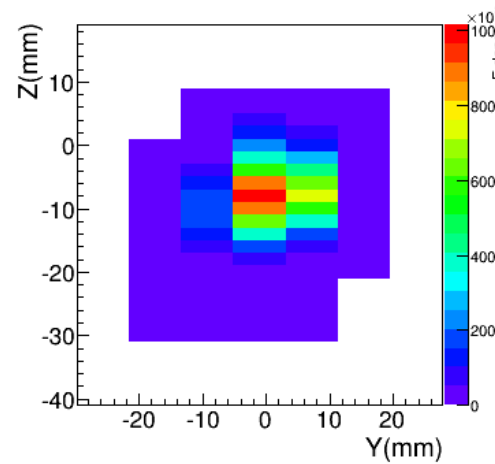
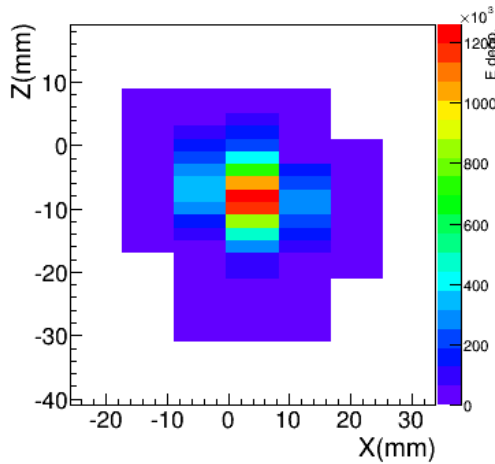
## diffusion after 1m drift





# simulation example 2

segmentation (7.5mm in x or y and 2mm in z)

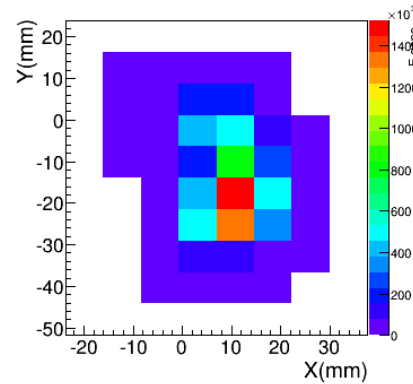
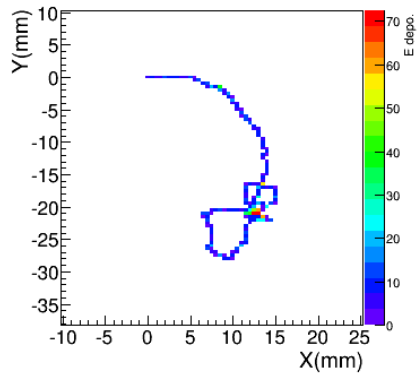
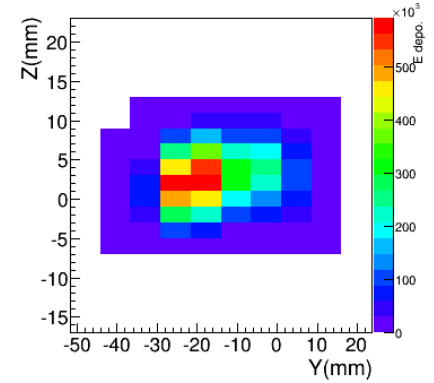
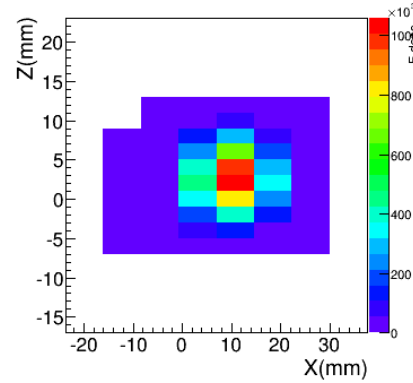
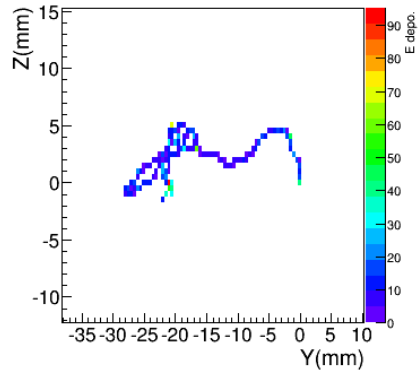
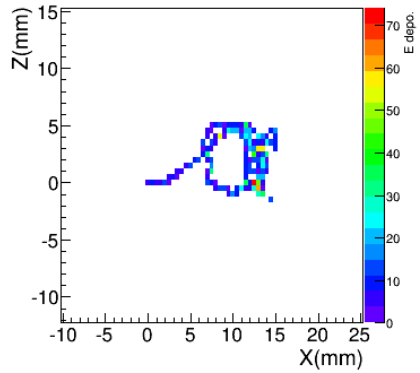


z marginalized by  
ELCC generation  
time(4 $\mu$ s)

# 2.5MeV electron example1

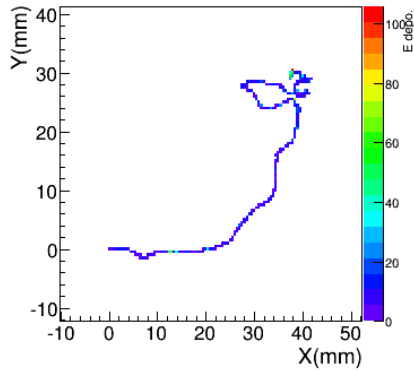
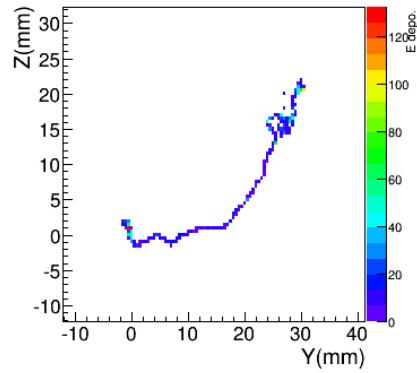
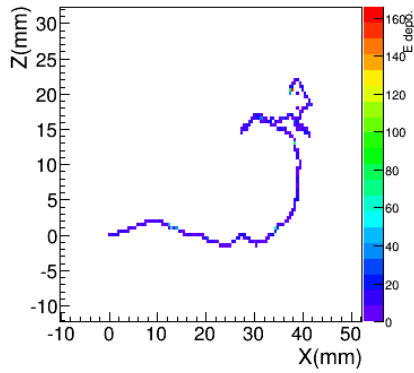
row track

detected track

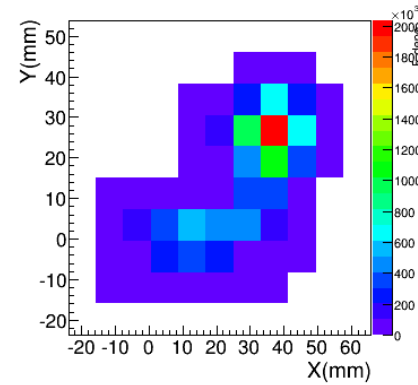
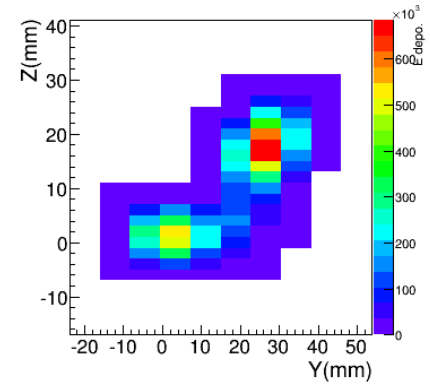
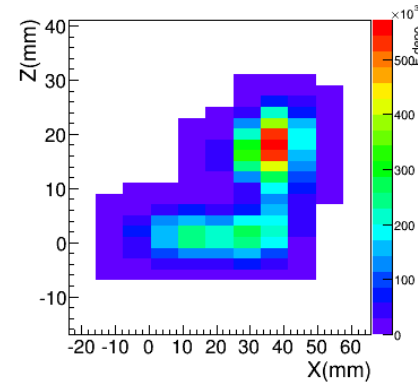


# 2.5MeV electron example2

row track

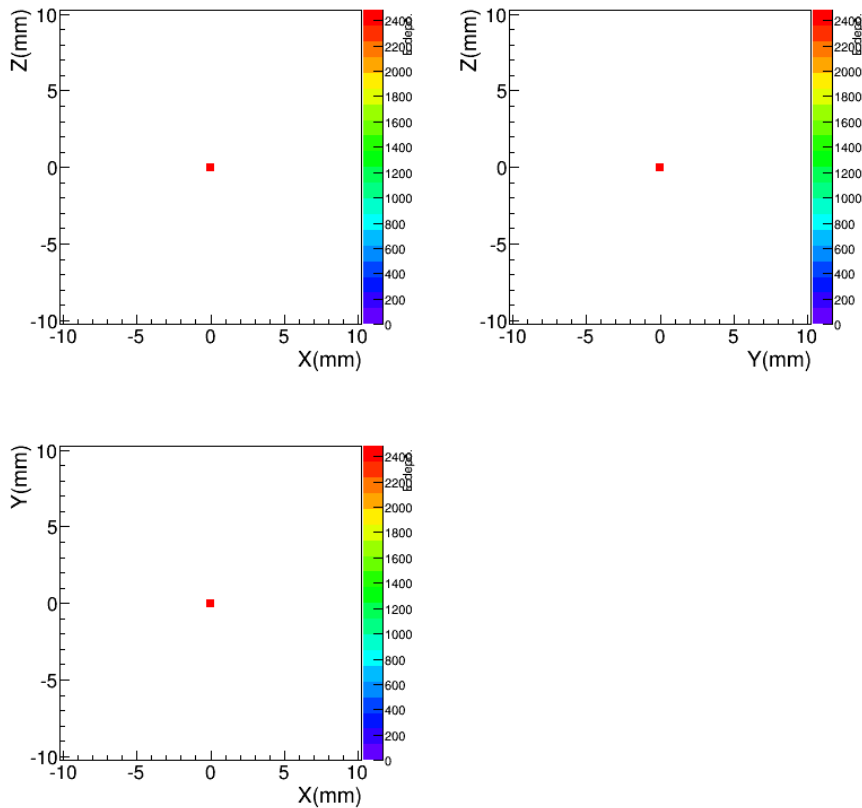


detected track

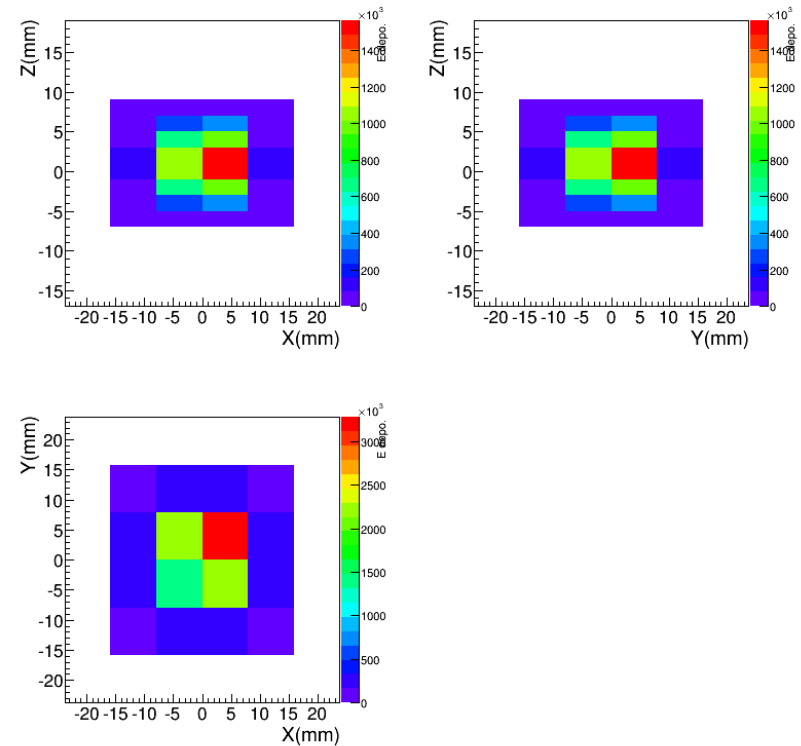


# 2.5MeV alpha example

row track



detected track



note that scale is very different from electron cases.

# trackイメージにより

- $\alpha$  backgroundは良く落とせそう
- $\gamma$  backgroundで、ちょうど2.5MeVで光電吸収したような事象は、すべて落とすことは難しい  
(コンプトン散乱等、複数の場所で電子を出すような事象は、落とせる)



## Radioactive contaminants in detector materials

After the decay of  $^{214}\text{Bi}$ , the daughter isotope,  $^{214}\text{Po}$ , emits a number of de-excitation gammas with energies above 2.3 MeV. The gamma line at 2447 keV, of intensity 1.57%, is very close to the  $Q$ -value of  $^{136}\text{Xe}$ . The gamma lines above  $Q_{\beta\beta}$  have low intensity and their contribution is negligible.

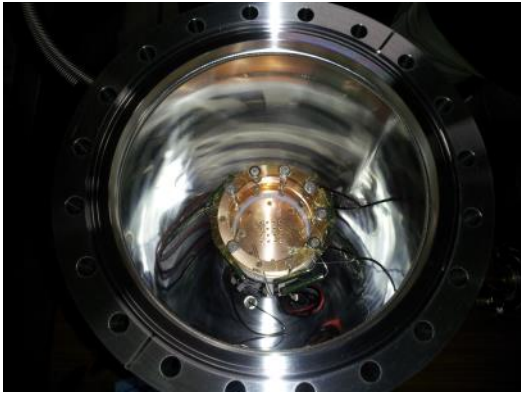
The daughter of  $^{208}\text{Tl}$ ,  $^{208}\text{Pb}$ , emits a de-excitation photon of 2614 keV with a 100% intensity. The Compton edge of this gamma is at 2382 keV, well below  $Q_{\beta\beta}$ . However, the scattered gamma can interact and produce other electron tracks close enough to the initial Compton electron so they are reconstructed as a single object falling in the energy region of interest (ROI). Photoelectric electrons are produced above the ROI but can lose energy via bremsstrahlung and populate the window, in case the emitted photons escape out of the detector. Pair-creation events are not able to produce single-track events in the ROI.

## Radon

Radon constitutes a dangerous source of background due to the radioactive isotopes  $^{222}\text{Rn}$  (half-life of 3.8 d) from the  $^{238}\text{U}$  chain and  $^{220}\text{Rn}$  (half-life of 55 s) from the  $^{232}\text{Th}$  chain. As a gas, it diffuses into the air and can enter the detector.  $^{214}\text{Bi}$  is a decay product of  $^{222}\text{Rn}$ , and  $^{208}\text{Tl}$  a decay product of  $^{220}\text{Rn}$ . In both cases, radon undergoes an alpha decay into polonium, producing a positively charged ion which is drifted towards the cathode by the electric field of the TPC. As a consequence,  $^{214}\text{Bi}$  and  $^{208}\text{Tl}$  contaminations can be assumed to be deposited on the cathode surface. Radon may be eliminated from the TPC gas mixture by recirculation through appropriate filters. There are also ways to suppress radon in the volume defined by the shielding. Radon control is a major task for a  $\beta\beta 0\nu$  experiment, and will be of uppermost importance for NEXT-100.

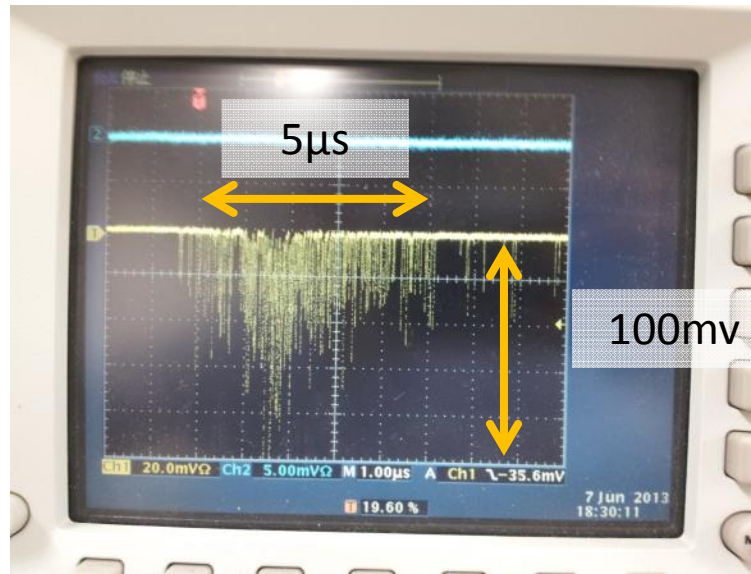
J.J.Gomez-Cadenas et.al, "Present status and future perspectives of the NEXT experiment", arXiv:1307.3914

# ELCC ver0



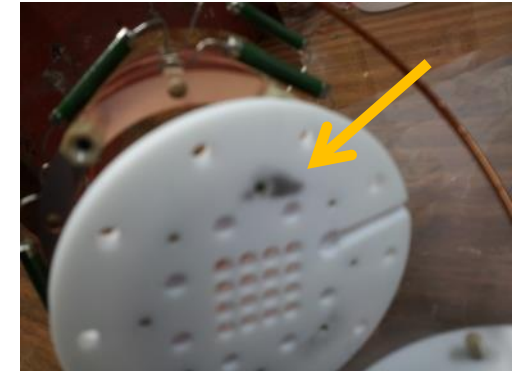
## ELCC ver0

読み出しには  
1 inch PMTにwave  
length shifter(TPB)  
を塗布して使用



## EL信号観測に成功！

光量が少ないため、PMTのパルス  
(~20ns)が重なって針のような信号  
に見えている。



ELCCでの放電



フィードスルーでの放電

## 問題点

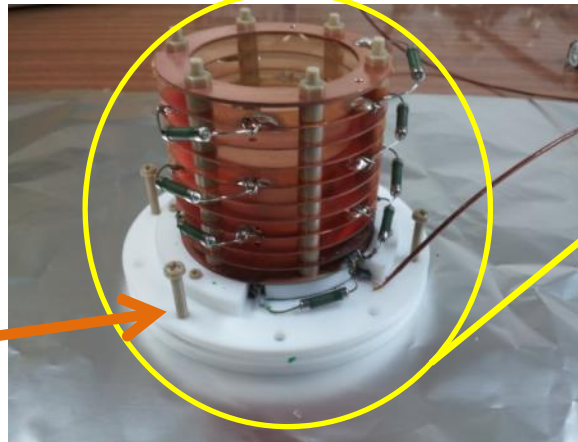
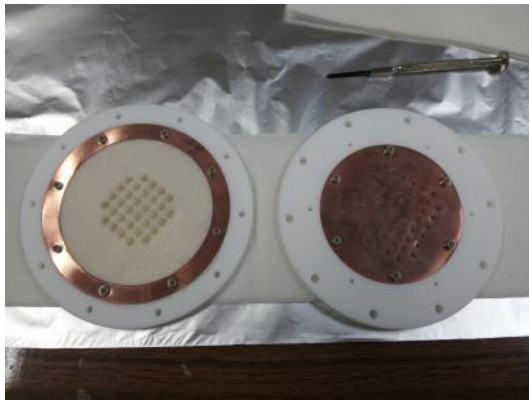
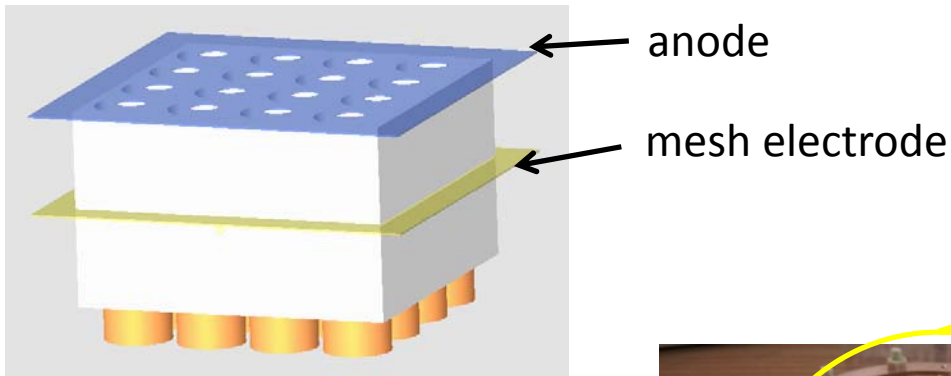
- フィードスルー、ELCCのギャップ間で放電
- 低電圧運用(ELCCのギャップ4mmで3kV)、電子の収集効率が悪いことによる光量不足

→ ver1で改善へ

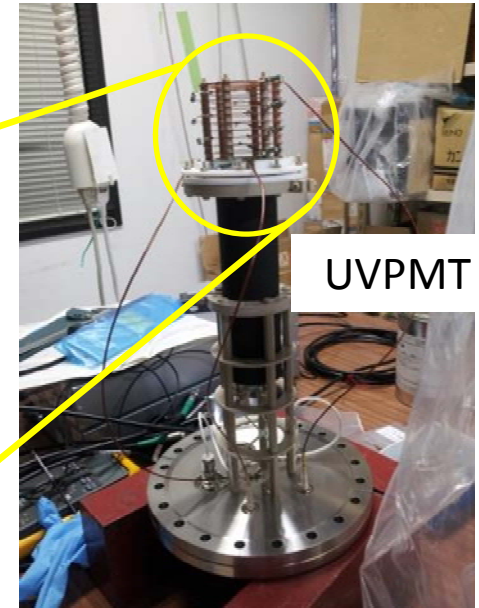


# ELCC ver1

- 読み出しにはUVPMT(浜ホト H3178)を用いた。

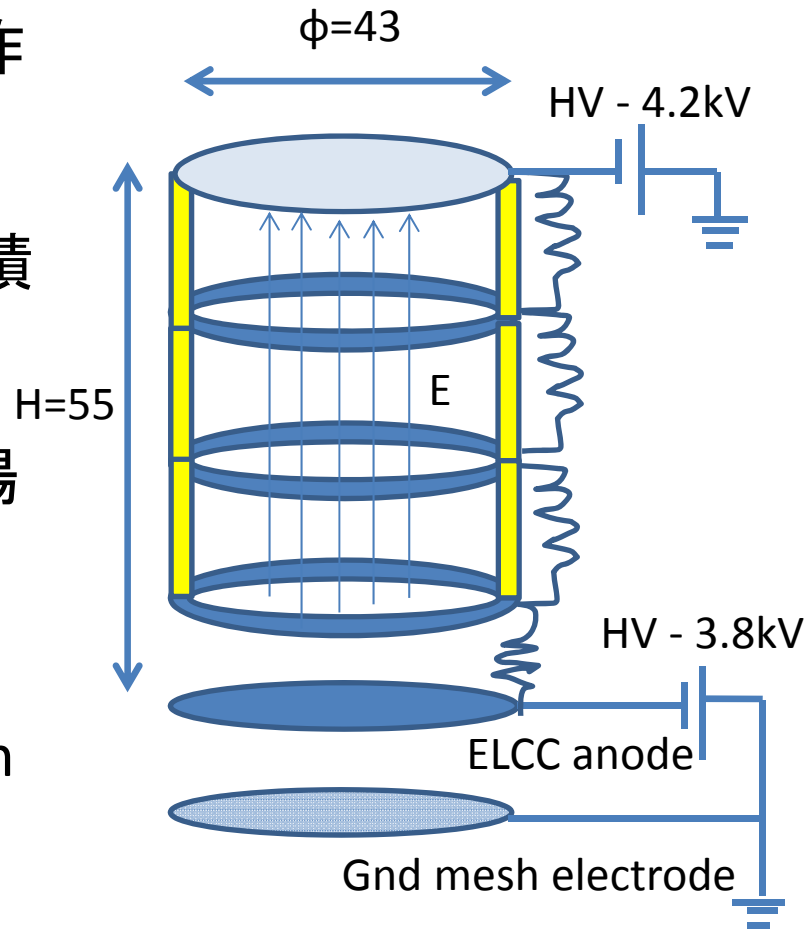


ELCCと  
フィールドケージ

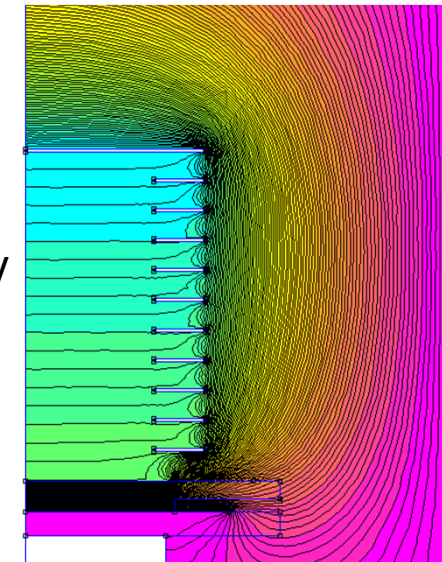


# フィールドケージ

- 電離した電子をELCCに導く電場を作る。
- 銅のリング(x11)を積み、抵抗分割する。
- z方向に一様な電場を構成する。
- 電子がdriftする領域は $\phi=43, H=55\text{mm}$
- 有効領域は $28 \times 28 \times 55\text{mm}$



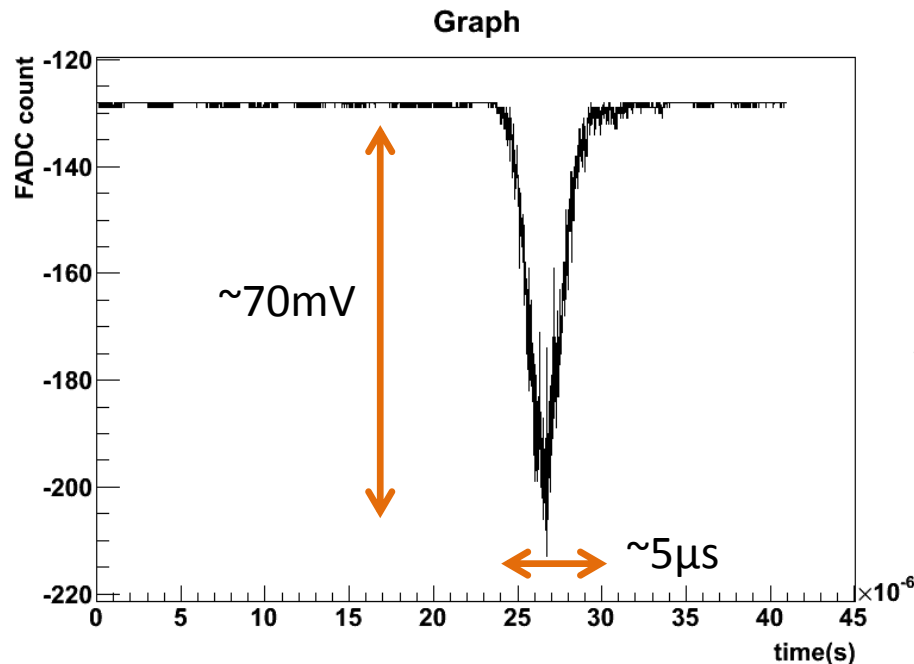
この部分



FEMMによる電場計算

# ELCC信号例

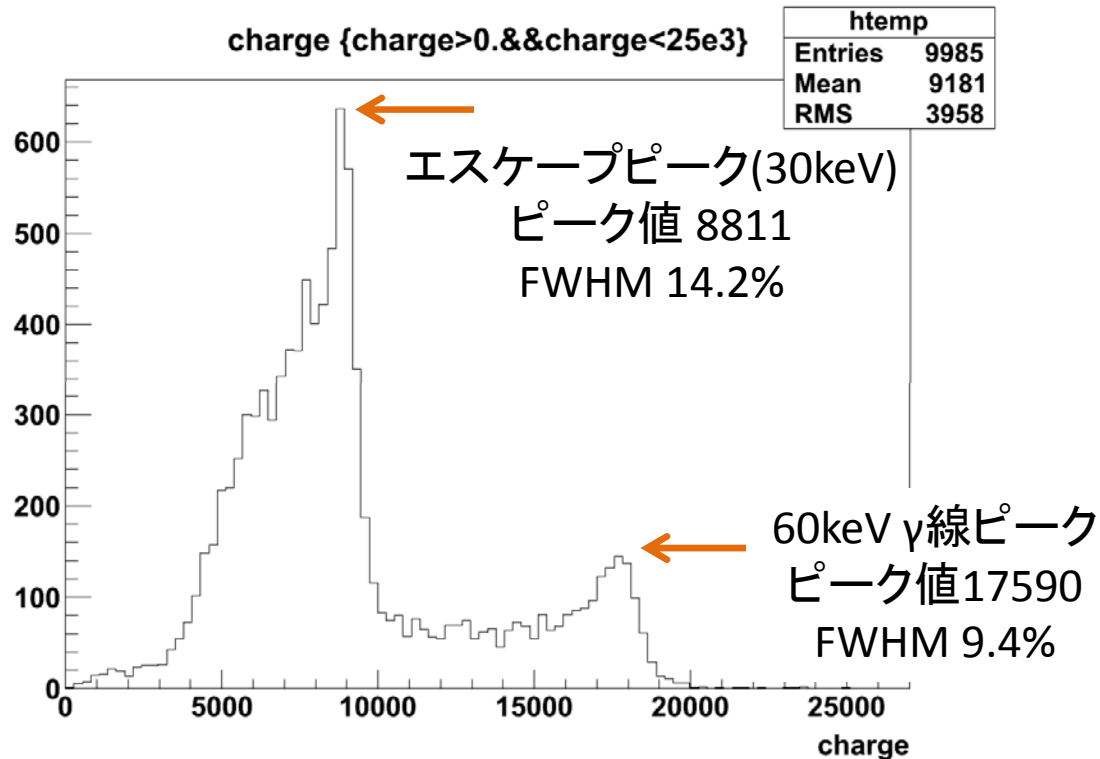
- 読み出しには100MHz FADC(CAEN V1724)を使用
- フィールドケージ内での電離電子のz方向の広がり( $\sim 1\text{cm}$ )に対して、ドリフト速度が $\sim 1 \times 10^5\text{cm/s}$ なので、信号幅は $\sim 10\mu\text{s}$ 程度と予想される。



60keV $\gamma$ 線の信号  
ver0に比べて、光量  
が増えている

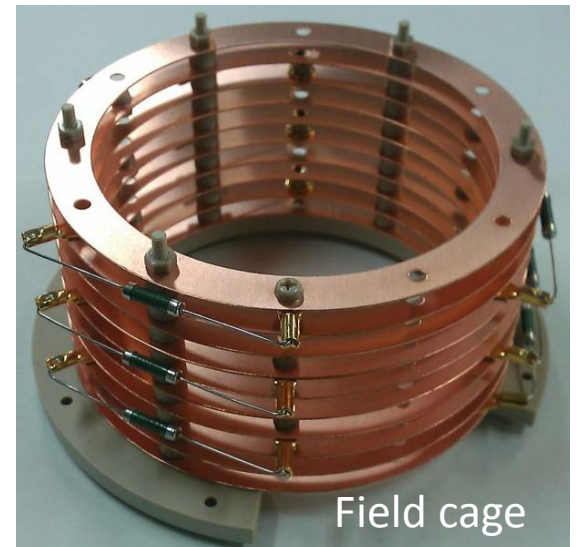
# 光量分布

- 2個のピークが見られる。-> 60keVピークとエスケープピーク(後述)
- 低値側に裾を引いているのはELCCの有効領域から漏れ出した電子の信号が失われたからであると考えられる。
- ピークの右側だけを用いてFWHMを概算評価した。



# Ver.3

- 1 MPa chamber
  - JIS配管
  - ガスケットにu-tightsealを採用 (helicoflexと同じようなメタルシール)
  - フィードスルーは、テフロンケーブルをエポキシモールド (TECSAM)
- MPPC plane (up to 64 channel)
- under construction



# Xeの純度

- PMT、ELCCからのアウトガス、フィードスルーのついたフランジからのリークにより、Xeの純度が悪化している可能性がある。
- 純度悪化による光量の減少
  - 電離電子の不純物ガスへの付着
  - Xe励起状態のクエンチ
$$\text{Xe}_2^* + \text{M} \rightarrow 2\text{Xe} + \text{M}$$
- 純度モニターの手法
  - Xe不純物によるdrift速度の変化(ELCC信号時間幅)を使う?

# Gas purification

MEG prototype

imeter) tungsten wires suspended inside the detector volume. Xenon is continuously evaporated, passed through an Oxisorb cartridge, a molecular sieve and a hot metal getter and condensed back in the detector. During several weeks of operation, the absorption length of scintillation light was increasing and has reached  $>150$  cm in four weeks. Mass-spectrometric analysis has shown the presence of water as a dominant impurity in xenon. Monte Carlo

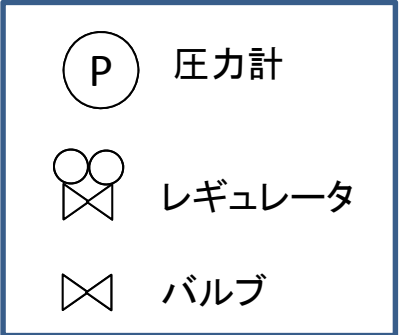
高圧のキセノンを循環できるか？ 純度を保つためには、金属ベローズポンプか。ただし、高圧のものがなかなか見つからない。循環させなくても、ゲッターやモレキュラーシーブをつないでおくだけではダメか？

# ガスの再利用

- キセノンが高いので、液体窒素で再収集することを考えている。
- が、常温に戻したときに10気圧以上になると高圧ガス製造装置になる可能性があり、ずっと冷やしていないといけない。。。
- お金があれば、冷凍庫を買う？ チェンバーをもう1個買う方が安い。

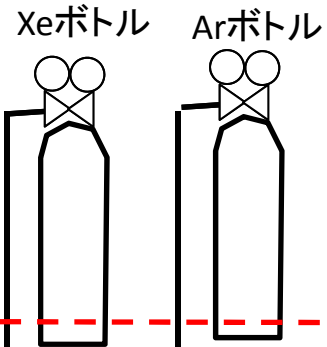


配管はSUS製とし、サイズは指定のない場合は1/4インチ  
 継手 指定のない場合はswagelok継手とする。  
 圧力逃し弁以外のバルブは、ダイヤフラムバルブとする。  
 ダイヤフラムバルブは特に指定のない場合は、DNシリーズとする。  
 清浄度は特に指定のない場合は、Swagelok SC-11仕様とする。

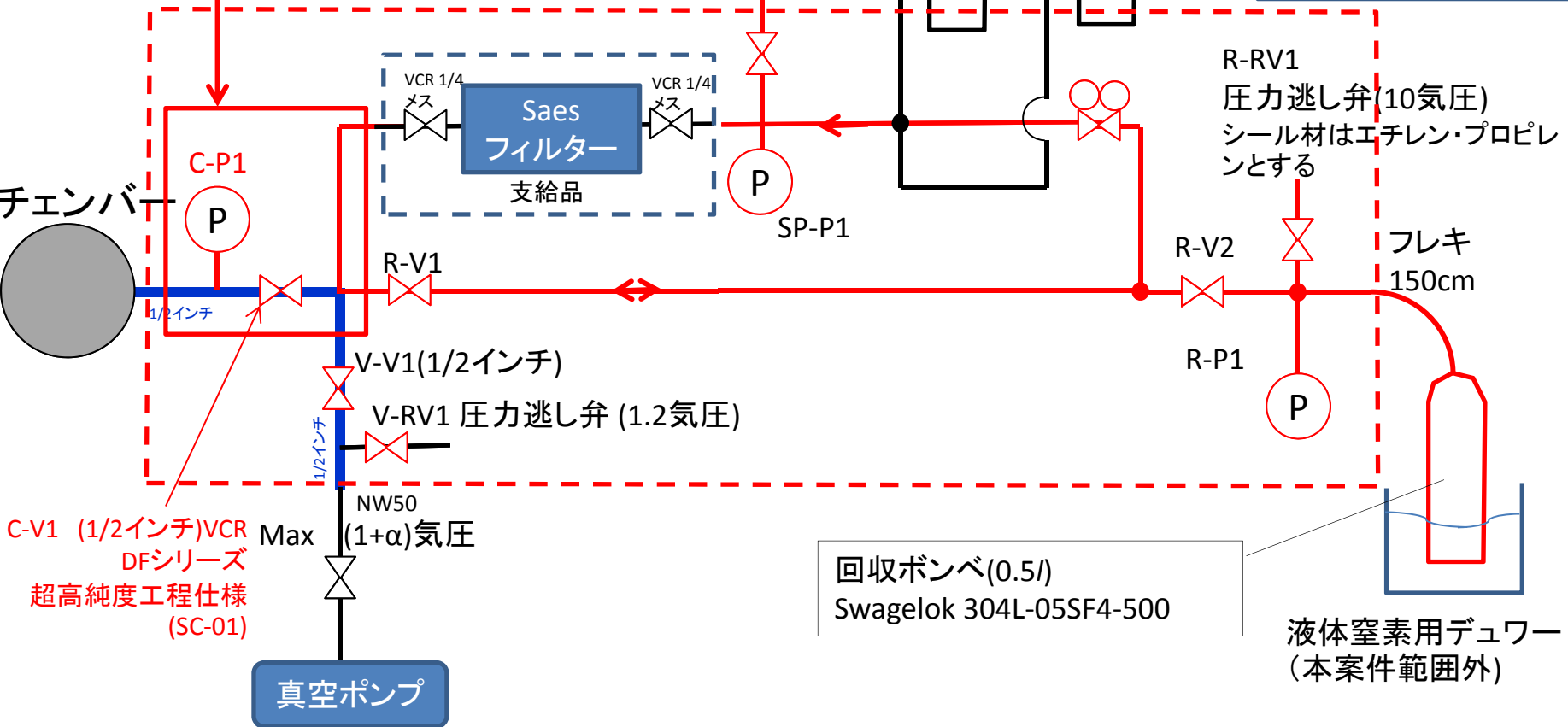


\*A ここだけ  
 クリーニングP VCR継手

SP-RV1  
 圧力逃し弁(10気圧)



R-RV1  
 圧力逃し弁(10気圧)  
 シール材はエチレン・プロピレンとする



C-V1 (1/2インチ)VCR  
 DFシリーズ  
 超高純度工程仕様  
 (SC-01)

Max  
 (1+α)気圧

回収ポンベ(0.5l)  
 Swagelok 304L-05SF4-500

液体窒素用デュワー  
 (本案件範囲外)

圧力計 : 圧力範囲は-1気圧から12気圧  
 バルブ : V-V1のみ流量調整(ベローズ)。その他は開閉(ダイヤフラム)。

# HV power supply

- 高電圧を容器の外からかけると、フィードスルーの耐圧が問題になるため、液体Arではコッククロフトウォルトン電源が検討されている。
  - 測定中は、AC入力を切る。電圧のモニターができない。Liq Arではノイズが問題だがEL読み出しでもだめか？
- 10気圧プロトタイプ
  - ドリフト電場 500V/cm
    - フィールドゲージはv2と同じ(ギャップ5.5mmx10段) → 275V/gap x 10段
  - EL電場 7.9~11kV @ 5mmギャップ
  - total max 13.75kV
  - これならば、コッククロフトウォルトンではなく、15kV耐圧のフィードスルーを2個使うべき。(+抵抗分割)
- 30気圧本番
  - ELCC電場 21kVまたはもっと。
  - ドリフト電場 1.5kV/cm。例えば、高

HOME 高耐圧絶縁トランス



ELTR-30K2型  
DC30kV重畳用の高耐圧絶縁トランスです。  
小型ION源重畳用です。

|     |                      |
|-----|----------------------|
| 電圧  | : 1φ 100:100 50/60Hz |
| 耐電圧 | : WV DC30kV          |
| 容量  | : 300VA              |
| 入力線 | : 3Pプラグ線 3m 1本       |
| 出力線 | : DC30kVシリコン線        |
| 出力部 | : 2連3PACコンセント 2個     |

# Direction sensitive dark matter search

- Extend the idea by D.R.Nygren, “Columnar recombination: a tool for nuclear recoil directional sensitivity in a xenon-based direct detection WIMP search”
- Our plan is to investigate the possibility of operation under magnetic field. → Enhance the recombination for the track parallel to the field
- To keep  $O(\text{eV})$  electron curvature  $< O(10)\mu\text{m}$ , hi field ( $>1\text{T}$ ) is necessary